



Two-dimensional simulations of strongly radiating plasmas with the RALEF-2D code

M. M. Basko

Institute for Theoretical and Experimental Physics, Moscow

Moscow, November 30, 2009

Equations of hydrodynamics



One fluid, one temperature, two spatial dimensions (either x,y, or r,z):

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0,$$

$$\frac{\partial (\rho \vec{u})}{\partial t} + \nabla \cdot (\rho \vec{u} \otimes \vec{u}) + \nabla p = 0,$$

$$\frac{\partial (\rho E)}{\partial t} + \nabla \cdot [(\rho E + p) \vec{u}] = \nabla \cdot (\kappa \nabla T) + Q_r + Q_{dep},$$

$$E = e + \frac{u^2}{2}, \quad e = e(\rho, T)$$

$\nabla \cdot (\kappa \nabla T)$ – energy deposition by thermal conduction (local), Q_r – energy deposition by radiation (non-local), Q_{dep} – eventual external heat sources.

Radiation transport



Transfer equation for radiation intensity I in the quasi-static approximation:

$$\cancel{\frac{1}{c} \frac{\partial I}{\partial t}} + \vec{\Omega} \cdot \nabla I = k_\nu (B_\nu - I), \quad I = I(t, \vec{x}, \nu, \vec{\Omega}), \quad B_\nu = B_\nu(\nu, T)$$

Quasi-static approximation: radiation transports energy infinitely fast (compared to the fluid motion) \Rightarrow the energy residing in radiation field at any given time is infinitely small !

Coupling with the fluid energy equation:

$$Q_r = -\nabla \cdot \left(\int d\nu \int \vec{\Omega} I d\vec{\Omega} \right) = \int_{4\pi} d\vec{\Omega} \int k_\nu (I - B_\nu) d\nu$$

Radiation transport adds 3 extra dimensions (two angles and the photon frequency) \Rightarrow **the 2D hydrodynamics becomes a 5D radiation hydrodynamics !**

Radiation transport in the diffusion approximation



In many cases the term “radiation hydrodynamics” is applied to hydrodynamic equations augmented with the multi-frequency diffusion equation

$$\frac{1}{c} \frac{\partial \mathcal{Q}_\nu}{\partial t} + \nabla \cdot \left(\mathcal{Q}_\nu \frac{\vec{u}}{c} \right) + \frac{1}{3} \mathcal{Q}_\nu \nabla \cdot \left(\frac{\vec{u}}{c} \right) = \nabla \cdot \left(\frac{1}{3k_\nu} \nabla \mathcal{Q}_\nu \right) + k_\nu \left(\frac{B_\nu}{c} - \mathcal{Q}_\nu \right)$$

for the spectral radiation energy density \mathcal{Q}_ν ; the radiation diffusion equation is coupled to the fluid energy equation via the energy exchange term

$$Q_r = \int k_\nu (c\mathcal{Q}_\nu - B_\nu) d\nu .$$

In this treatment, information about the angular dependence of the radiation field is lost (e.g. the CHIC code, P.-H. Maire, J. Breil), and the 2D hydrodynamics becomes a 3D radiation hydrodynamics.

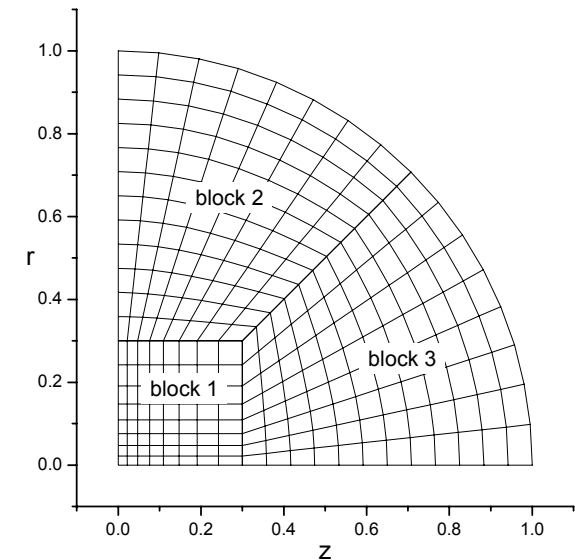
Whereas finite-difference treatment of the multi-frequency diffusion is similar to that of thermal conduction, solution of the transport equation requires a completely different numerical scheme.

Numerical scheme for hydrodynamics



To solve numerically the equations of hydrodynamics in two dimensions, we employ the CAVEAT-2D (LANL, 1990) hydrodynamics package. The CAVEAT numerical scheme has the following properties:

- ❖ it uses cell-centered principal variables on a multi-block structured quadrilateral mesh (either in the x-y or r-z geometry);
- ❖ is fully conservative and belongs to the class of second-order Godunov schemes;
- ❖ the mesh is adapted to the hydrodynamic flow by applying the ALE (arbitrary Lagrangian-Eulerian) technique;
- ❖ the numerical method is based on a fast non-iterative Riemann solver (J.K.Dukowicz, 1985), easily applied to arbitrary equation of state.



Numerical scheme for thermal conduction



The key ingredient to the RALEF-2D code is the **SSI (symmetric semi-implicit)** method of E.Livne & A.Glasner (1985), used to incorporate **thermal conduction** and **radiation transport** into the 2D Godunov method.

The numerical scheme for thermal conduction (M.Basko, J.Maruhn & A.Tauschwitz, J.Com.Phys., **228**, 2175, 2009) has the following features:

- ❖ it uses cell-centered temperatures from the FVD (finite volume discretization) hydrodynamics on distorted quadrilateral grids,
- ❖ is fully conservative (based on intercellular fluxes with an SSI energy correction for the next time step),
- ❖ (almost) unconditionally stable,
- ❖ space second-order accurate on all grids for smooth κ ,
- ❖ symmetric on a local 9-point stencil,
- ❖ computationally efficient.

Numerical scheme for radiation transport



Our numerical scheme for radiation transport (not published yet) has the following basic properties:

- ❖ radiation coupling to the fluid is combined with thermal conduction within the unified **SSI approach**;
- ❖ the angular dependence is treated by applying the classical **S_n method** with $n(n+2)$ fixed photon propagation directions over the 4π solid angle;
- ❖ the scheme is **non-conservative** in the sense that energy deposition by radiation Q_r is calculated not via elementary fluxes,
- ❖ the algorithm has the important property of correct transition to **the diffusion limit** when mesh cells become optically thick.

Presently, we do not see any practical alternative to the SSI method to cope with numerical instabilities in cases where radiation transport has a strong dynamic impact on hydrodynamics.

The method of short characteristics



For each angular direction $\vec{\Omega}_L$ and frequency ν , the radiation field I is found by solving the transfer equation

$$\vec{\Omega}_L \cdot \nabla I = k_\nu (B_\nu - I), \quad I = I(t, \vec{x}, \nu, \vec{\Omega}_L),$$

with the **method of short characteristics** (A.Dedner, P.Vollmöller, JCP, **178**, 263, 2002). **Mesh nodes** are chosen as collocation points for the radiation intensity I .

Advantages:

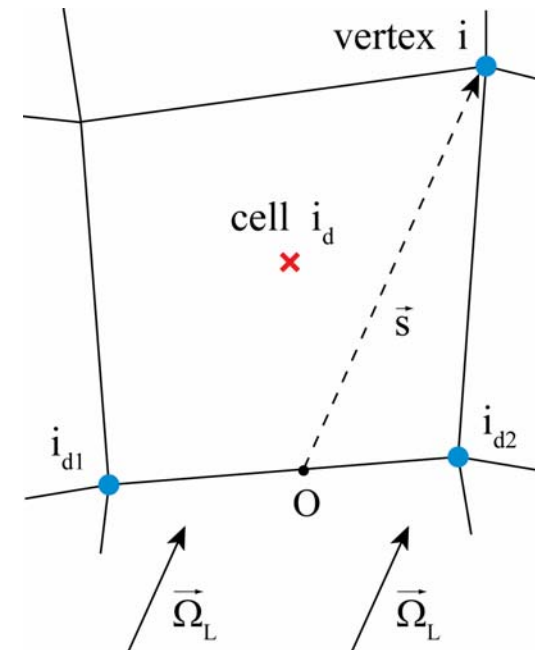
- ❖ even on strongly distorted meshes, it is guaranteed that light rays pass through each mesh cell;
- ❖ the algorithm is generally computationally more efficient than that of long characteristics.

Disadvantages:

- ❖ a significant amount of numerical diffusion in space.

Difference with R.Ramis: in the MULTI-2D code I is assigned to cell edges.

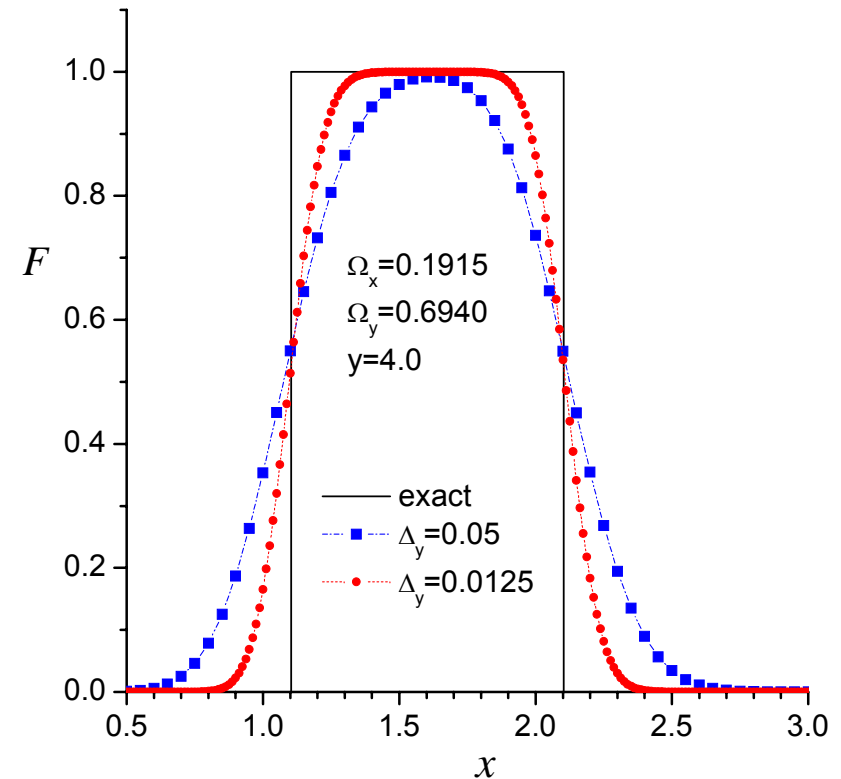
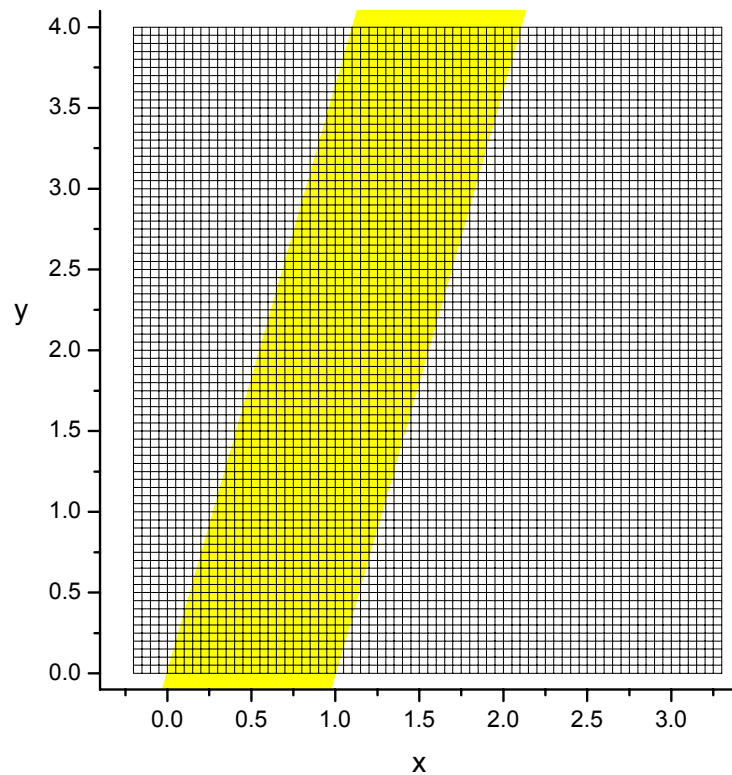
Only Cartesian geometry with straight characteristics has been treated so far.



Numerical diffusion: a searchlight beam in vacuum



The method of short characteristics produces a significant amount of numerical diffusion for light beams with sharp edges.

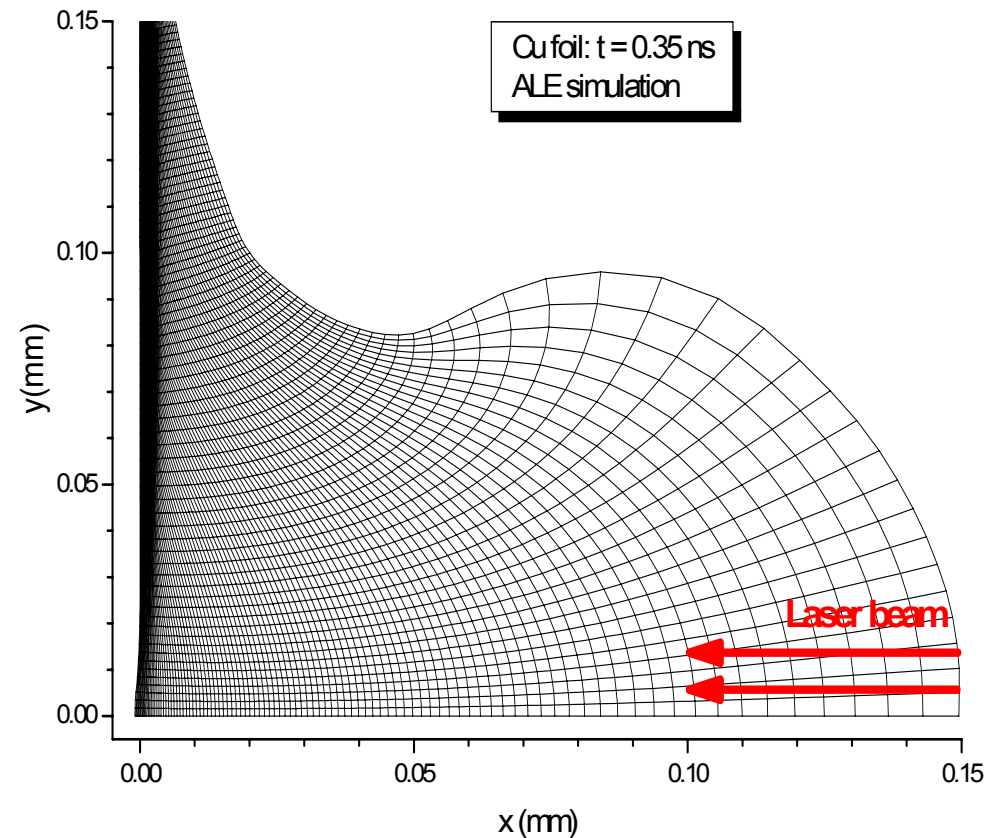
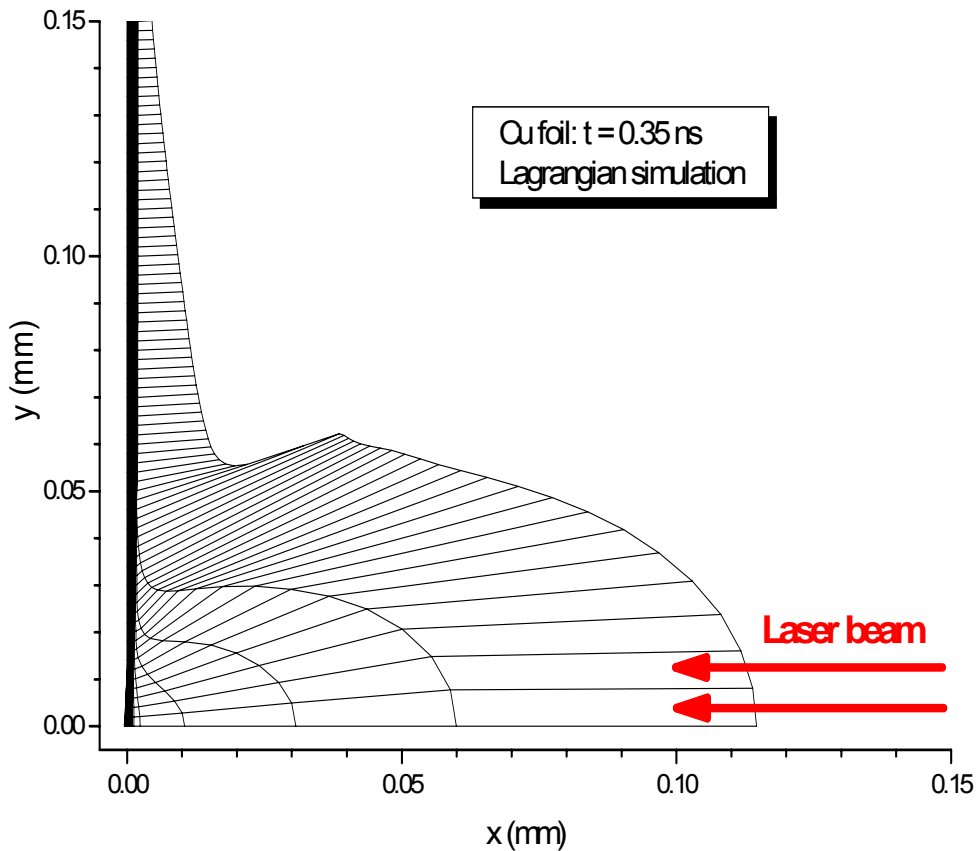


However, for thermal radiation a certain amount of numerical diffusion may be more an advantage than a shortcoming.

Vital importance of the ALE technique



The problem: a thin foil is irradiated by an intense laser beam.

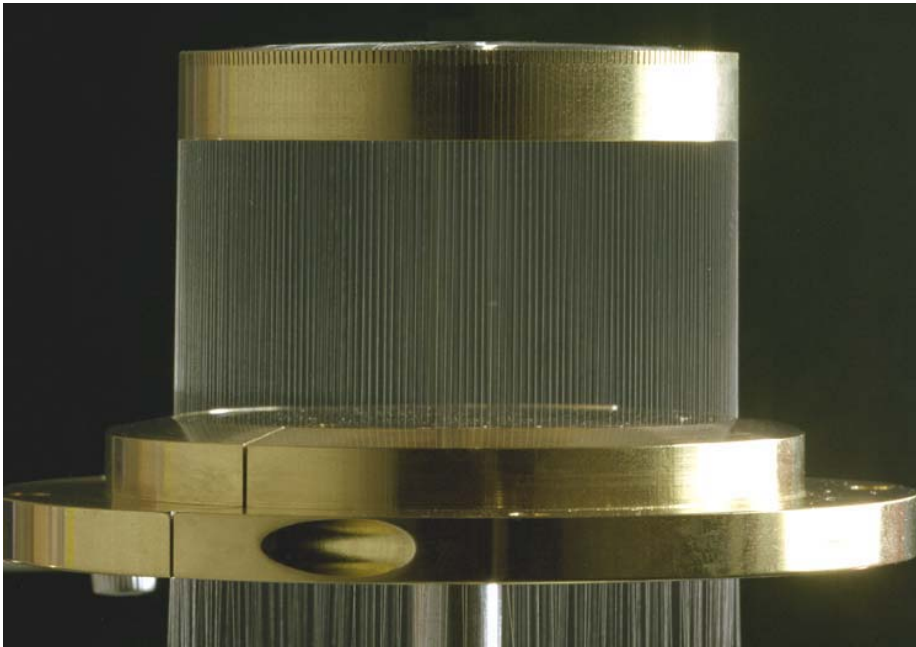


**Problem 1: A strongly radiating central Z-pinch
in tungsten multi-wire arrays**

Multi-wire Z-pinches (Sandia, Angara-5)



40-mm diameter array of 240, 7.5- μm -diam. wires.



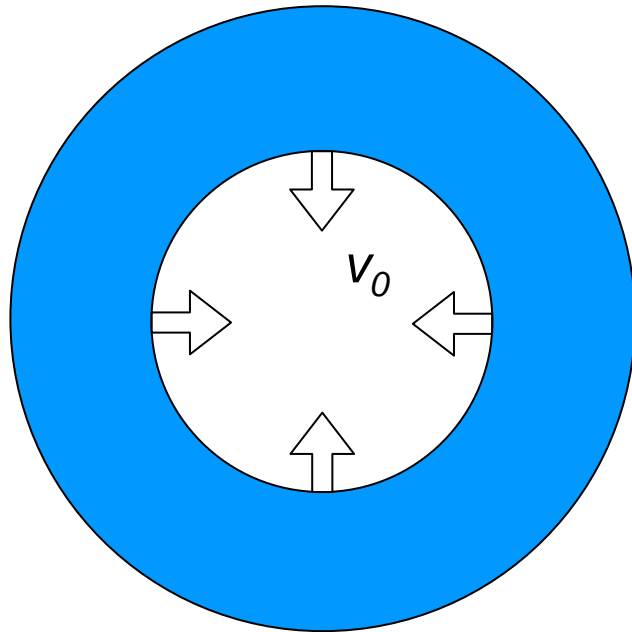
Z-machine at Sandia (USA):

- 11.5 MJ stored energy
- 19 MA peak load current
- 40 TW electrical power to load
- 100-250 TW x-ray power
- 1-1.8 MJ x-ray energy

Problem statement for the Angara-5 parameters



Cylindrical implosion of an initially cold tungsten plasma cloud



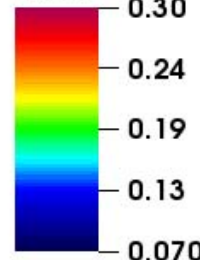
Initial shell parameters:

- radial thickness 2 mm;
- implosion velocity $v_0 = 400$ km/s;
- uniform temperature $T_0 = 20$ eV;
- mass per unit length 0.286 mg/cm;
- initial kinetic energy 23 kJ/cm;
- far from the axis, mass is uniformly distributed over the radius;
- possible influence of magnetic field is neglected.

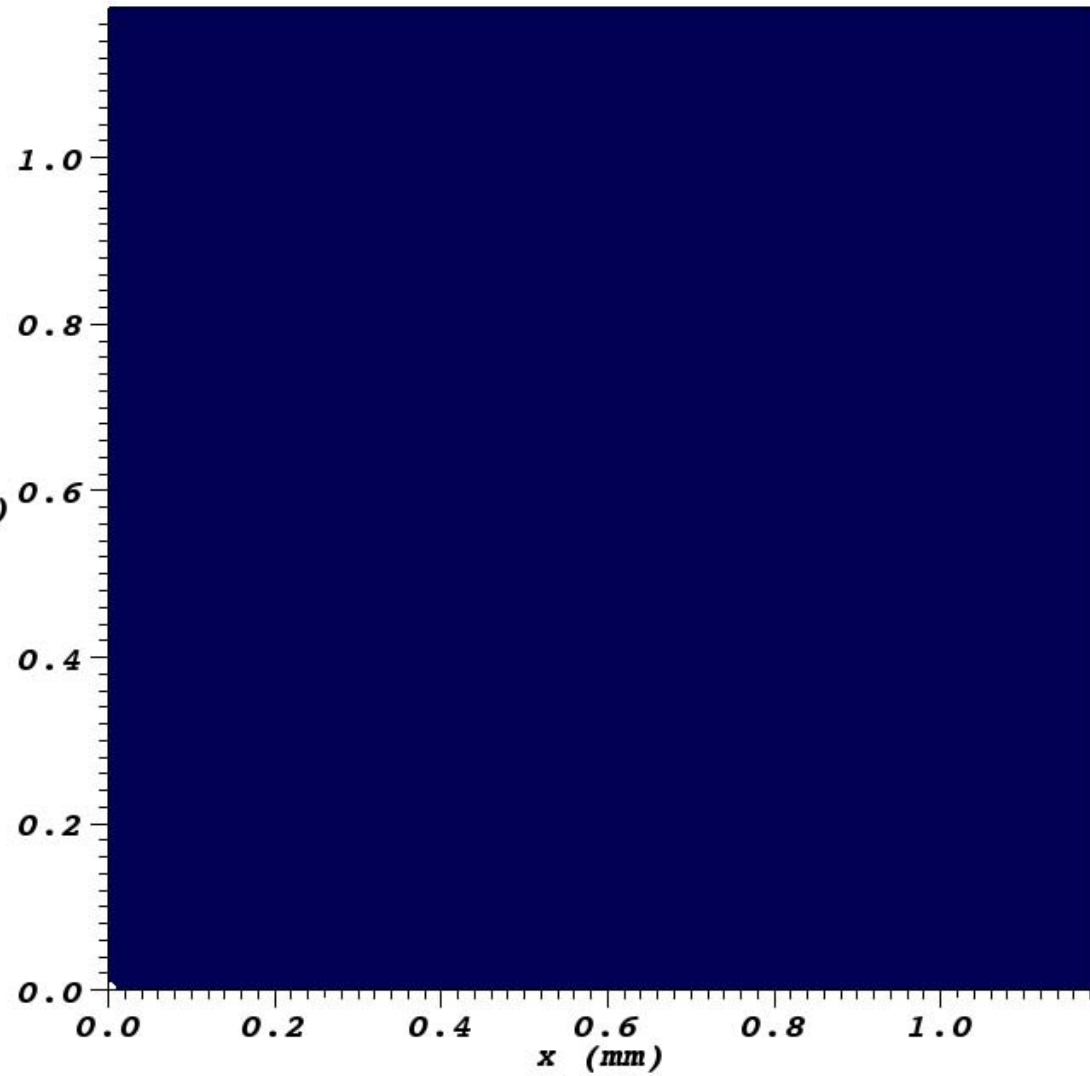
In its present formulation, the problem is one-dimensional.

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Pseudocolor
Var: T

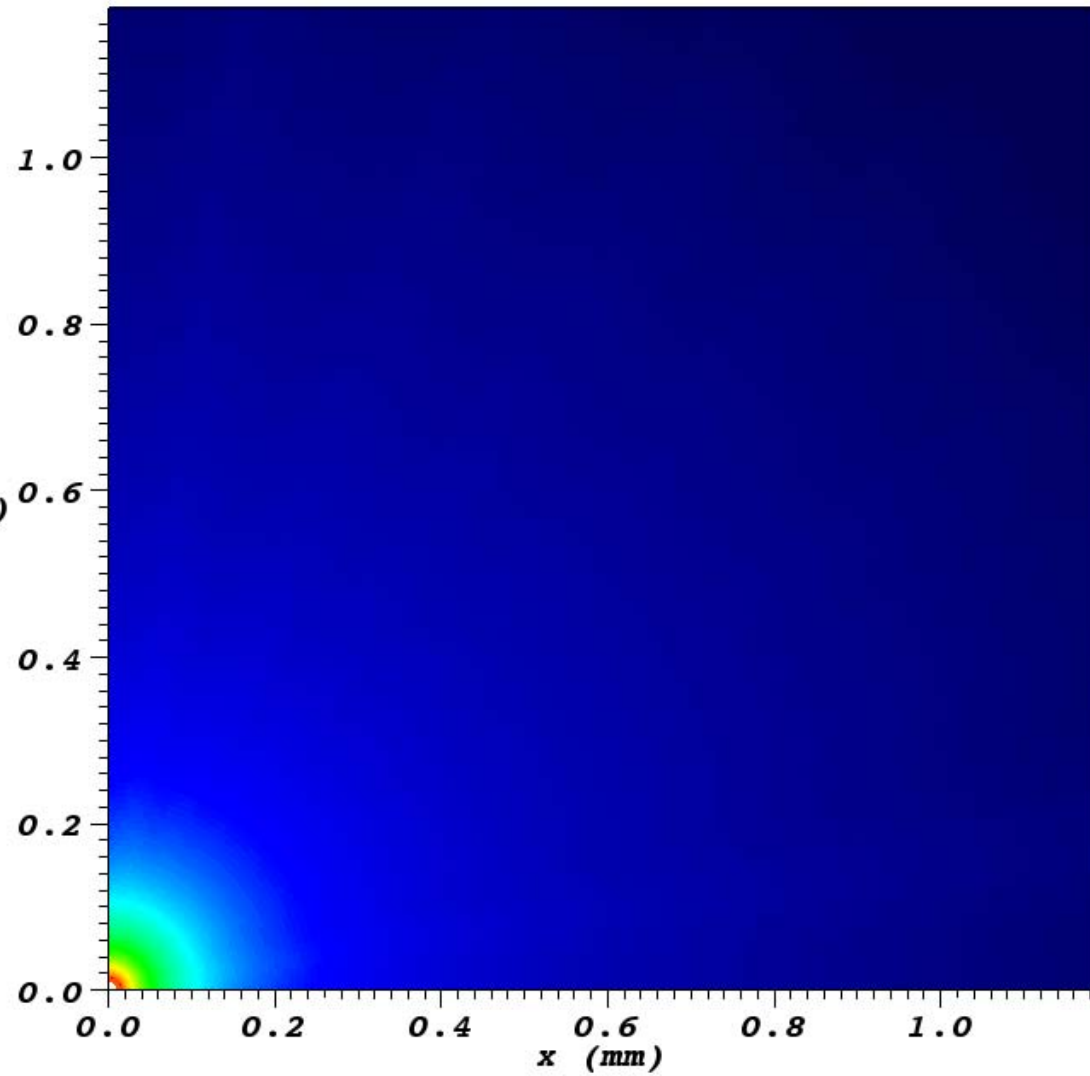
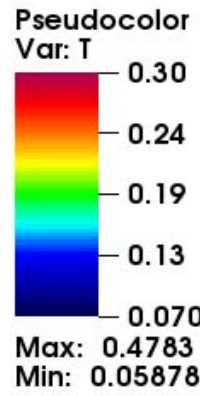


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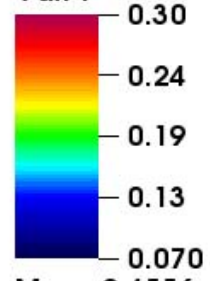
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Cycle: 1



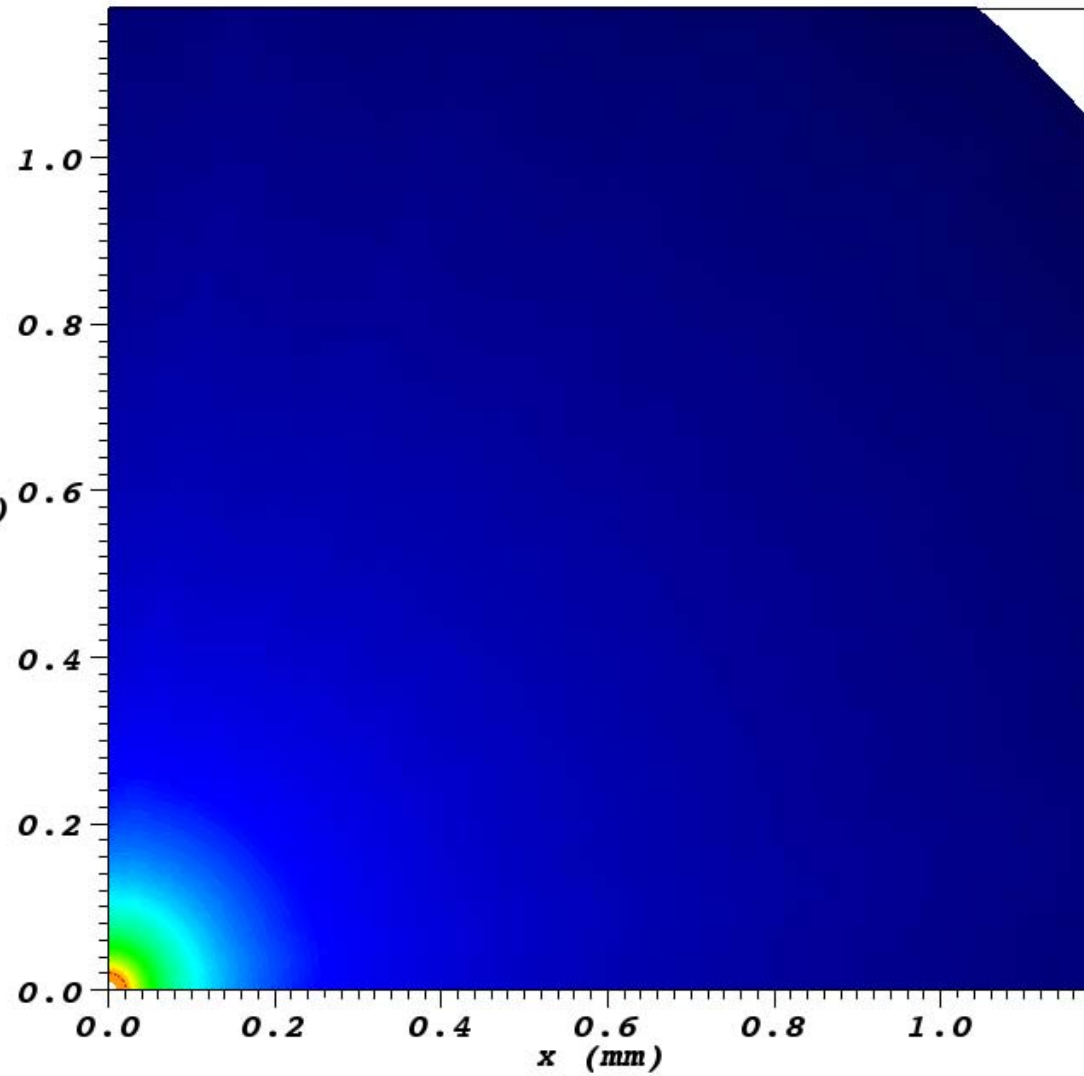
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DB: vtkall002.vtk
Cycle: 2

Pseudocolor
Var: T

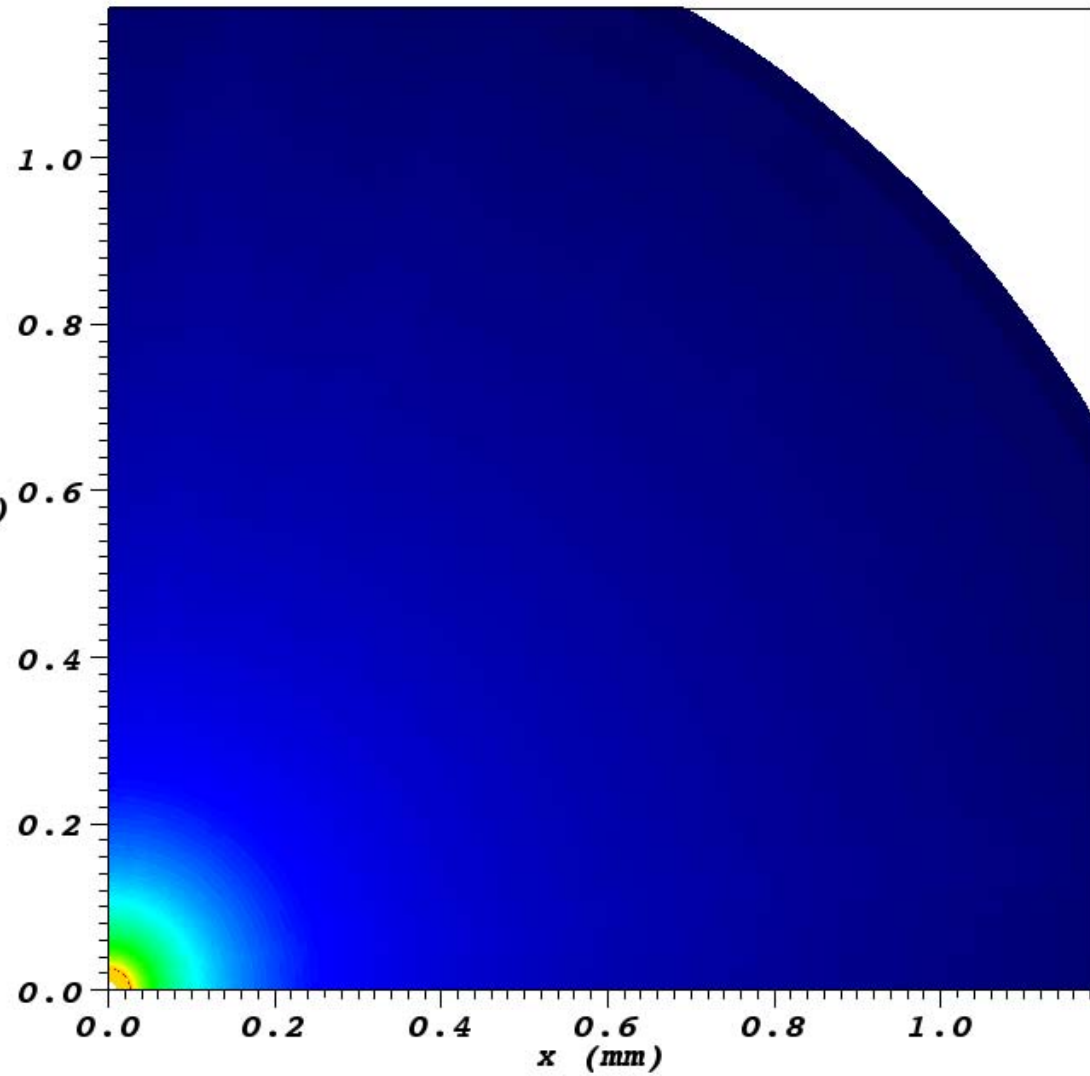
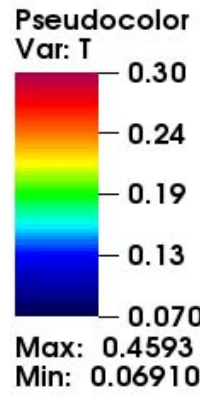


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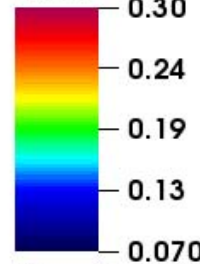
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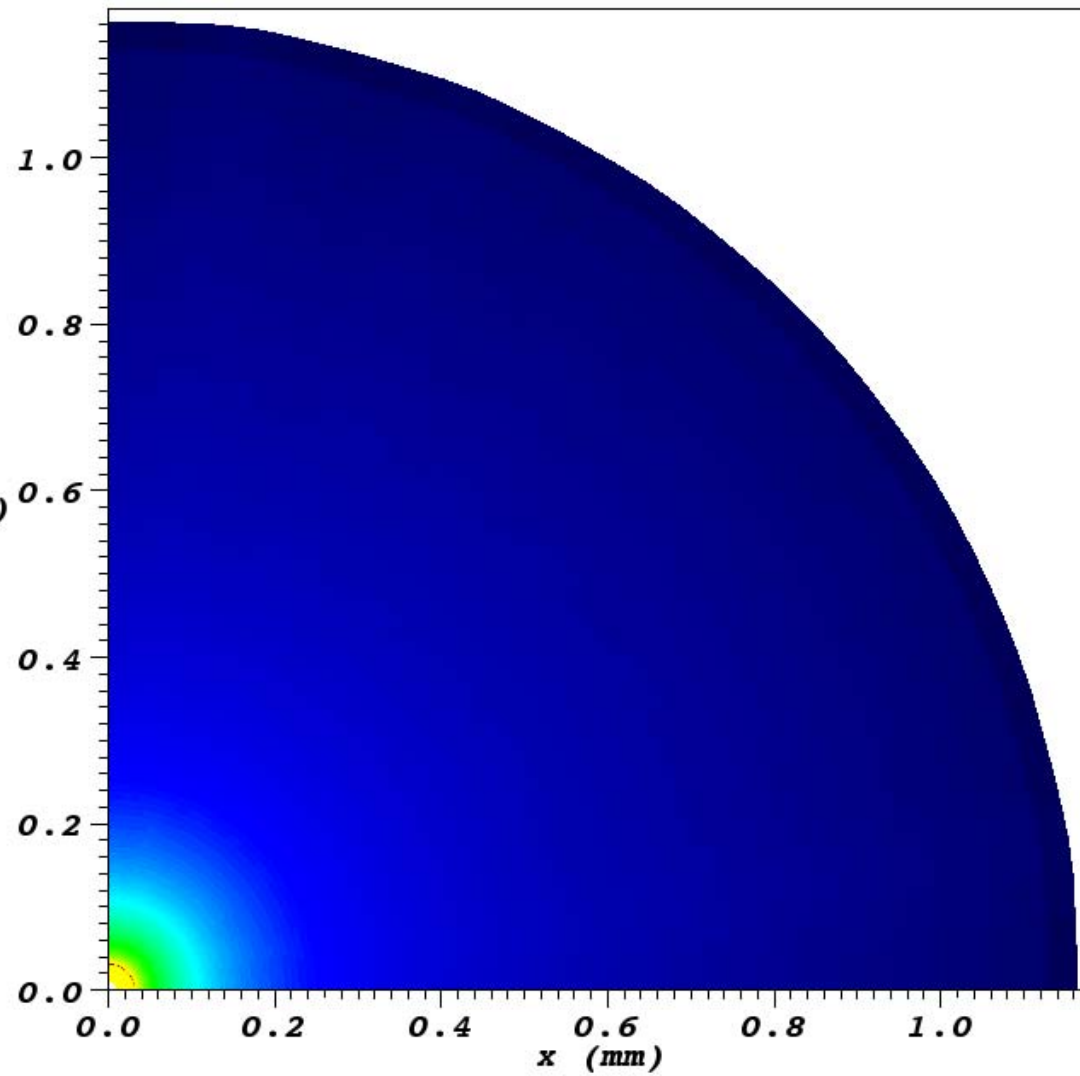
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Cycle: 4

Pseudocolor
Var: T



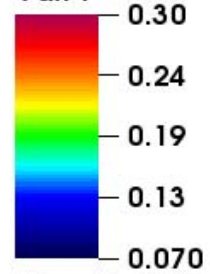
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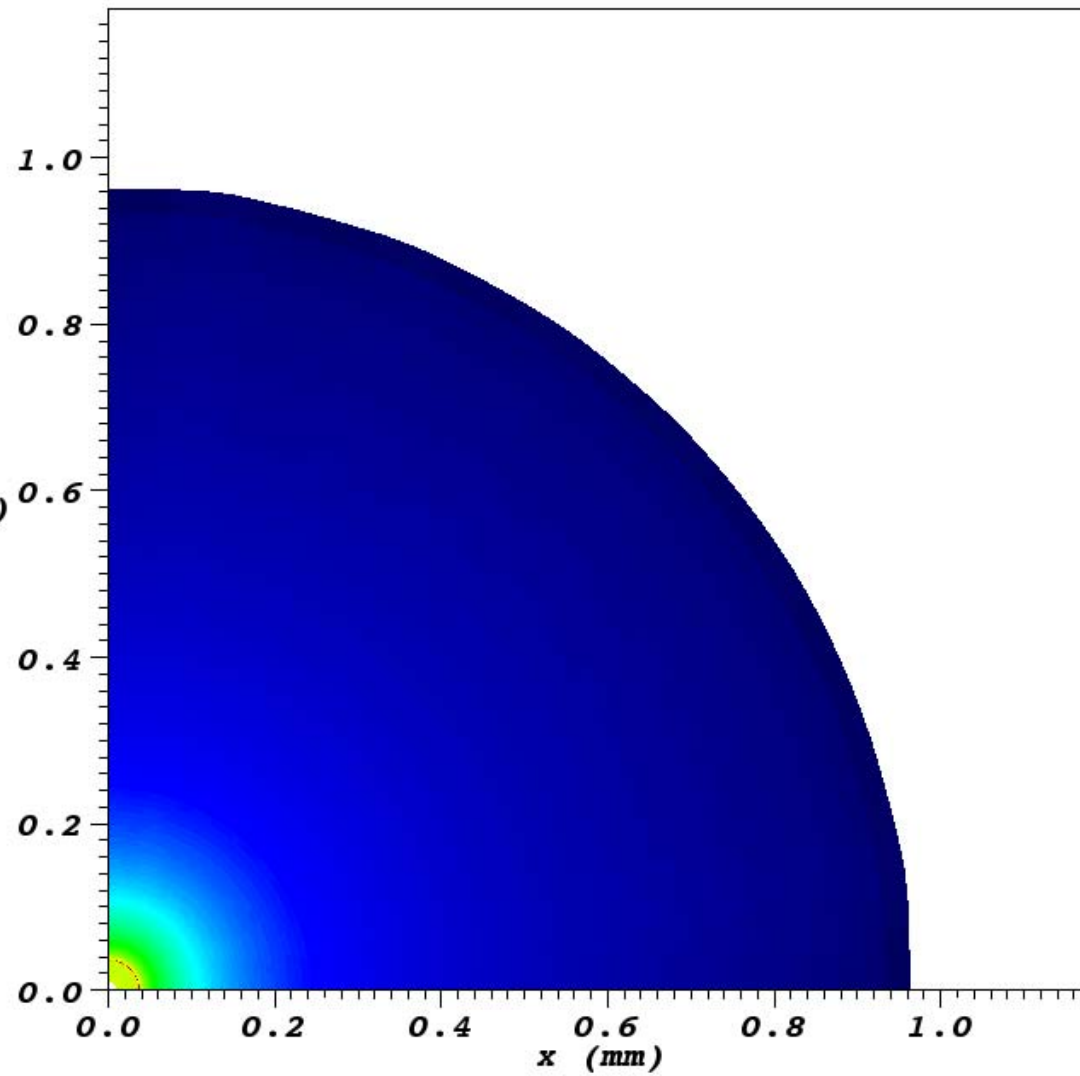
DB: vtkall005.vtk
Cycle: 5

Pseudocolor
Var: T



Max: 0.4352
Min: 0.07398

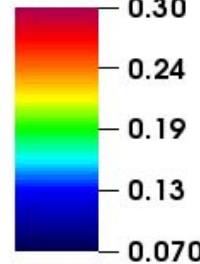
y (mm)



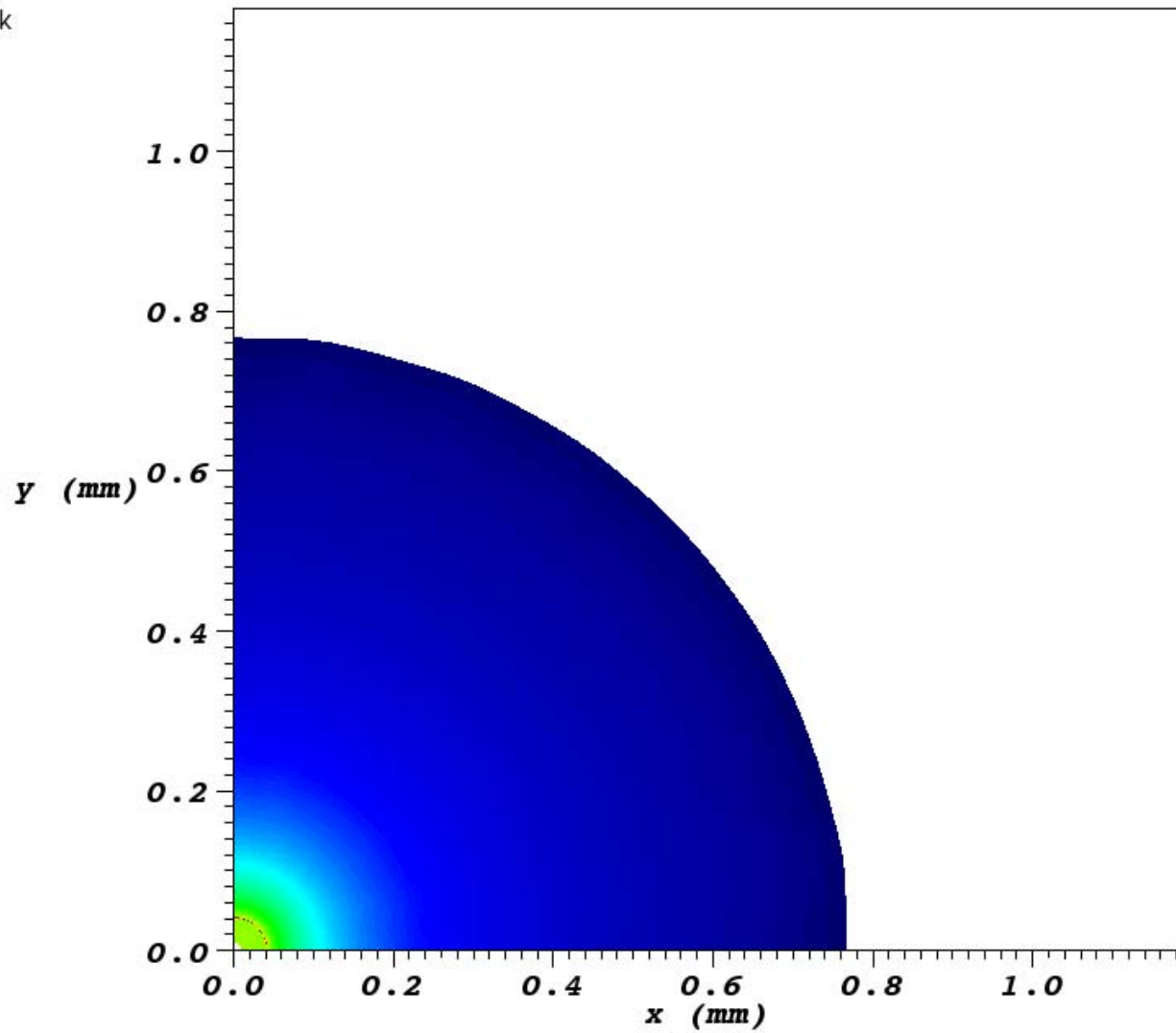
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DB: vtkall006.vtk
Cycle: 6

Pseudocolor
Var: T



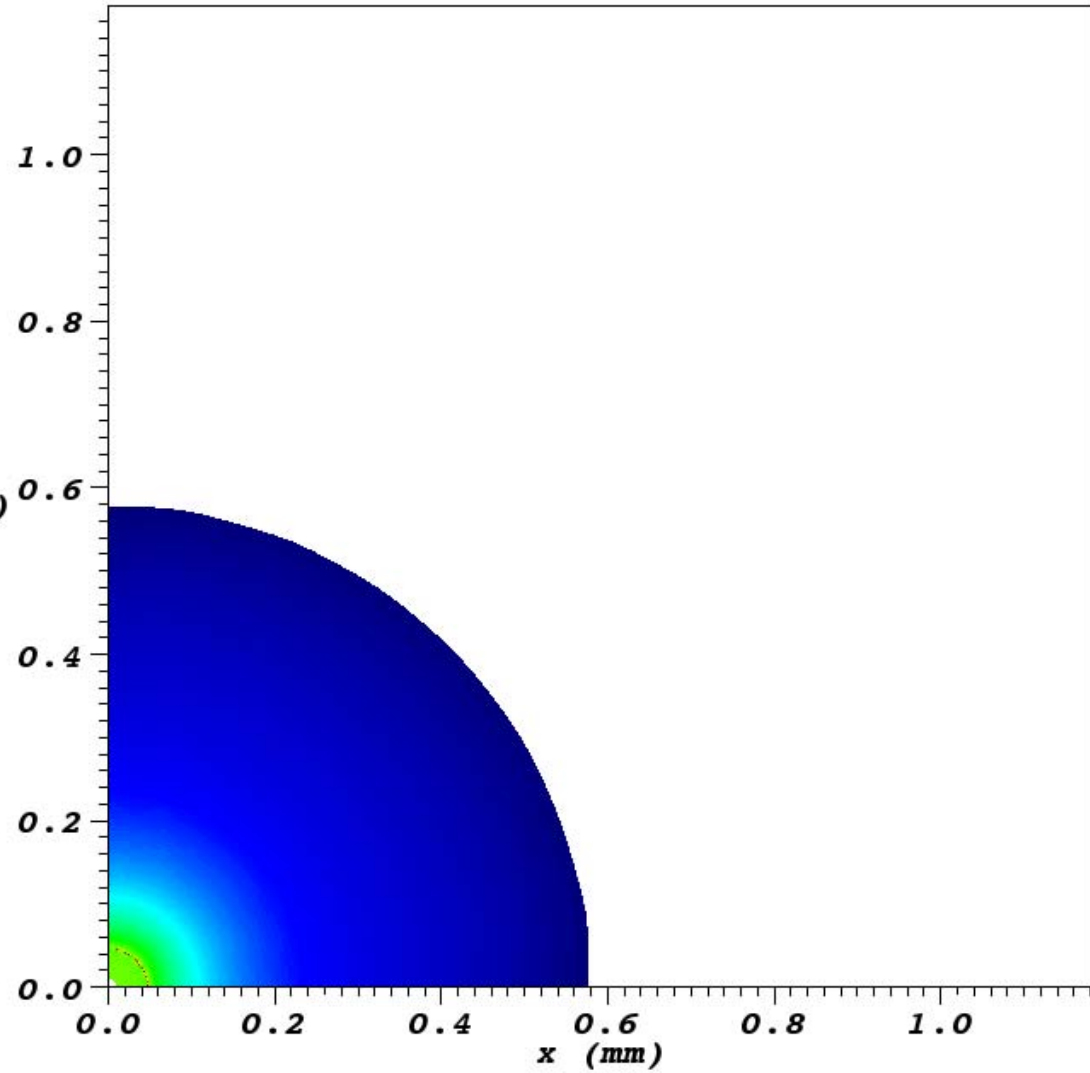
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Min: 0.07787



user: Mikhail Basko
Sat Nov 07 12:14:22 2009

DB: vtkall007.vtk
Cycle: 7

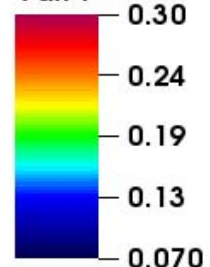
Pseudocolor
Var: T
0.30
0.24
0.19
0.13
0.070
Max: 0.4798
Min: 0.08286



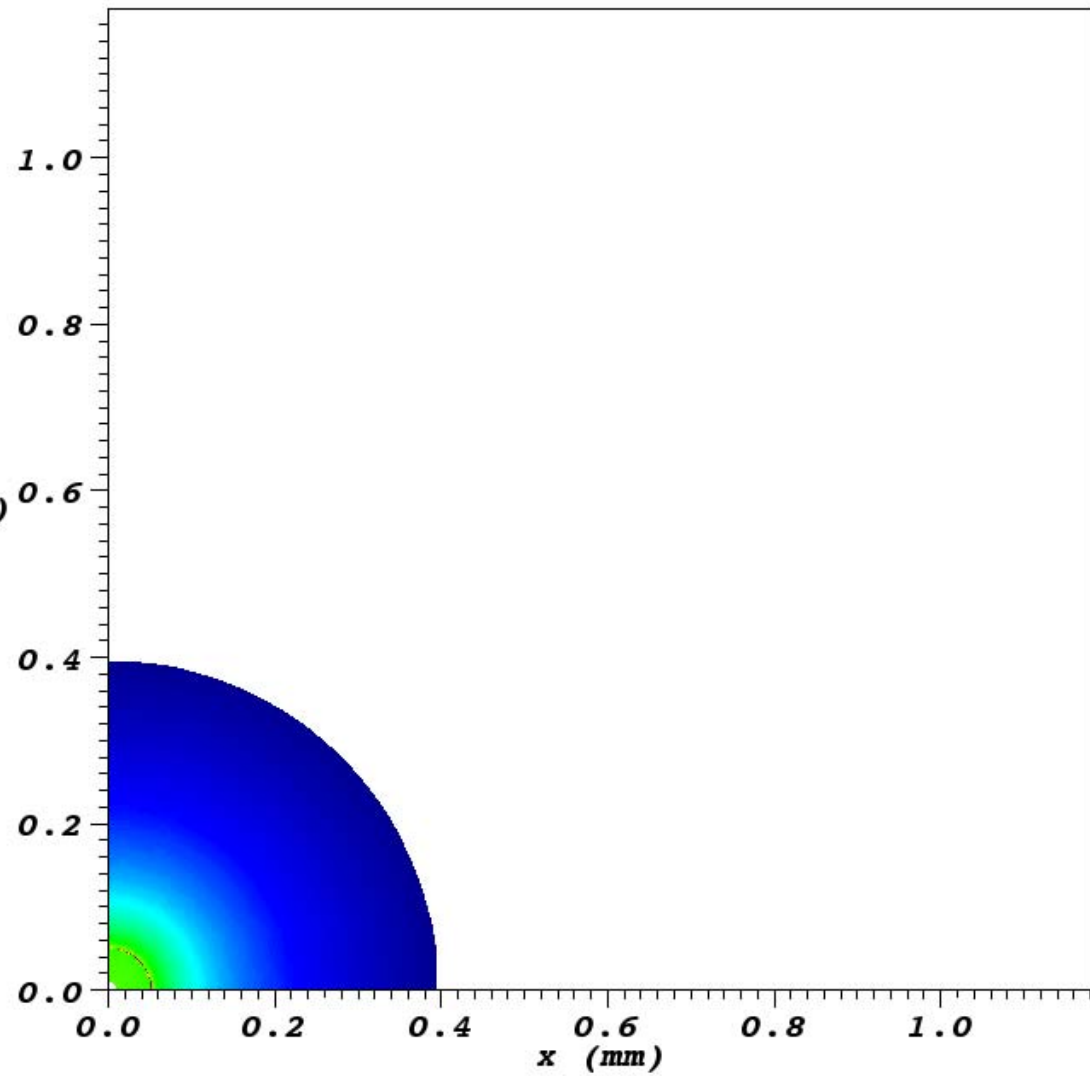
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DB: vtkall008.vtk
Cycle: 8

Pseudocolor
Var: T



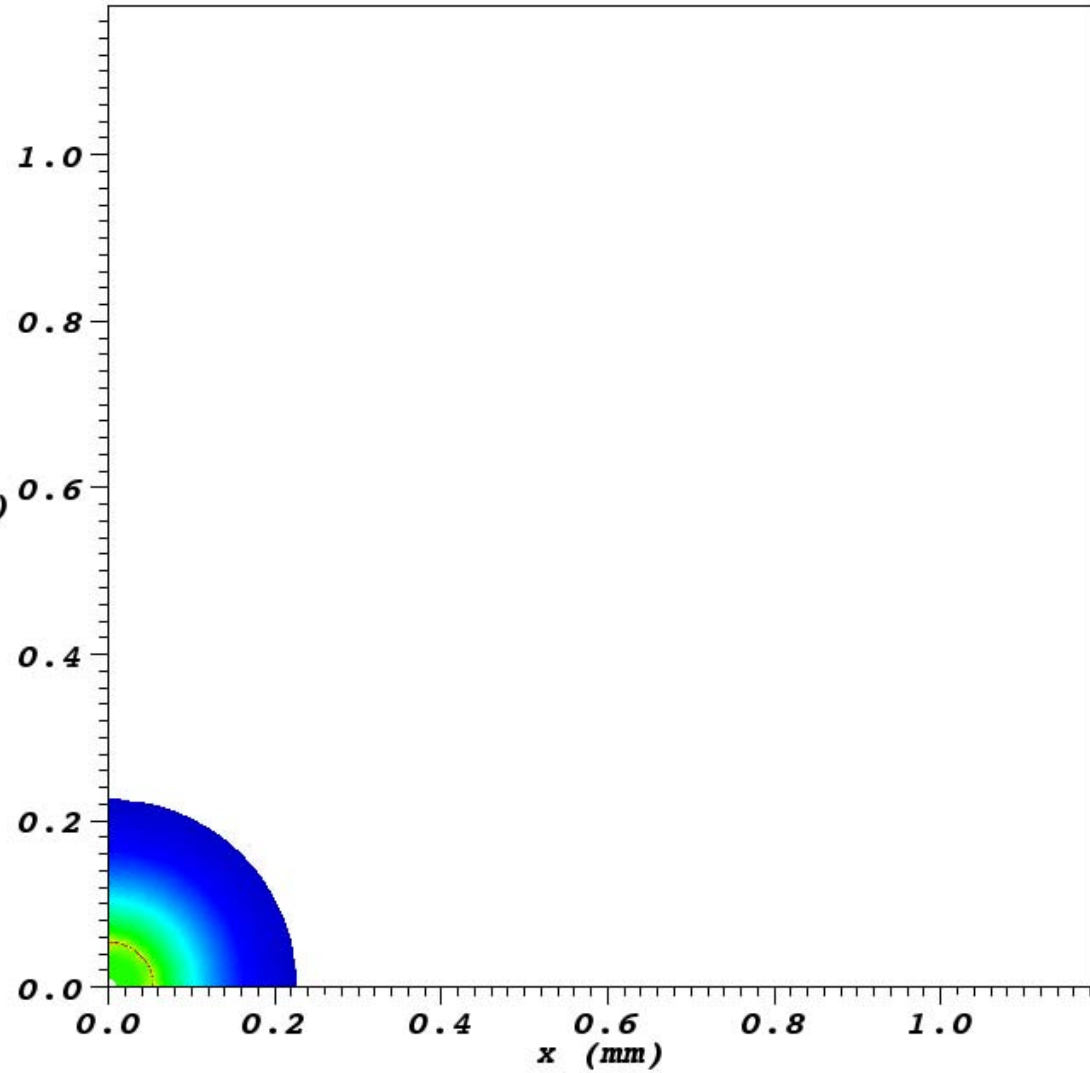
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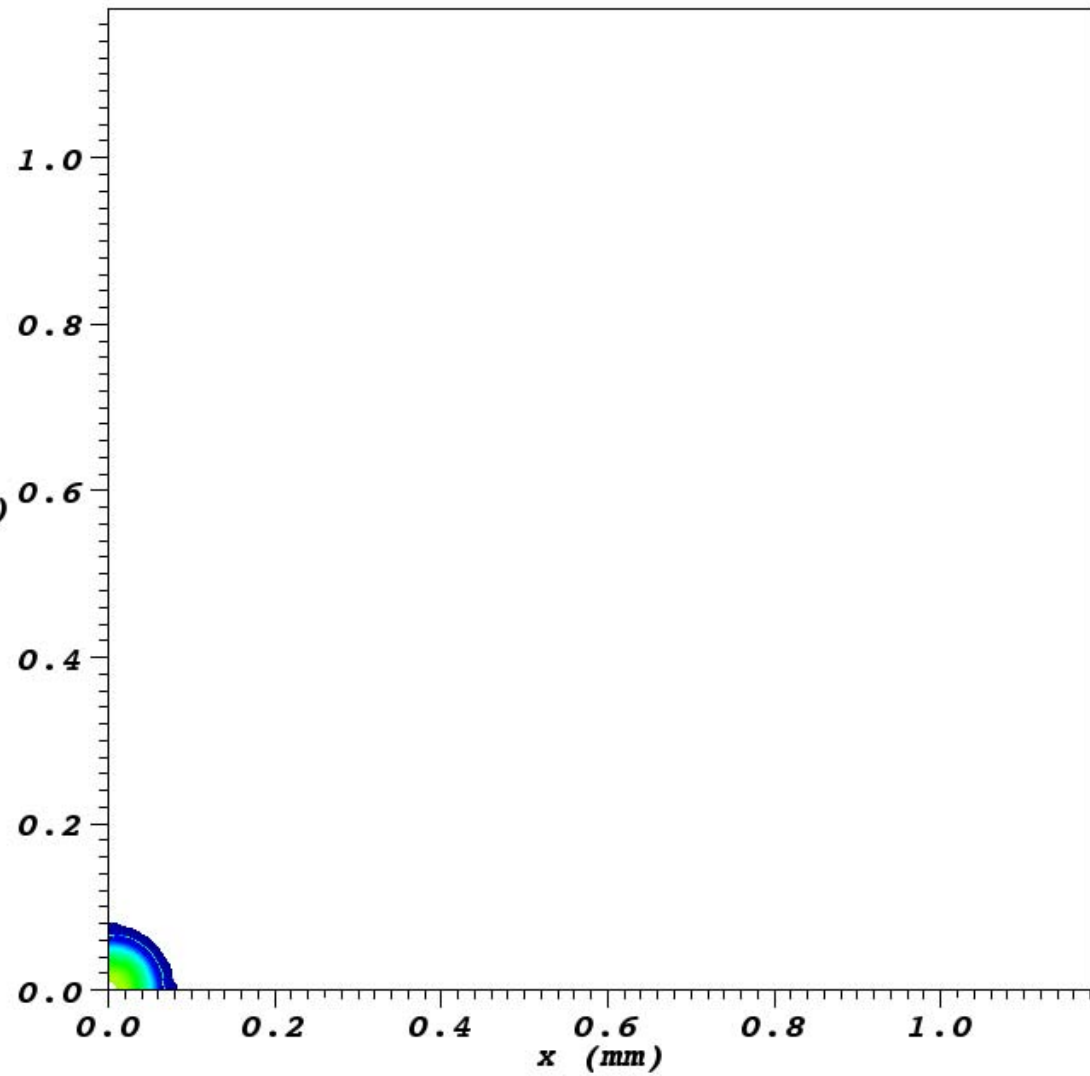
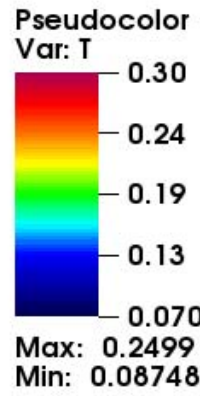
DB: vtkall009.vtk
Cycle: 9

Pseudocolor
Var: T
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0.24
0.19
0.13
0.070
Max: 0.5017
Min: 0.1090



user: Mikhail Basko
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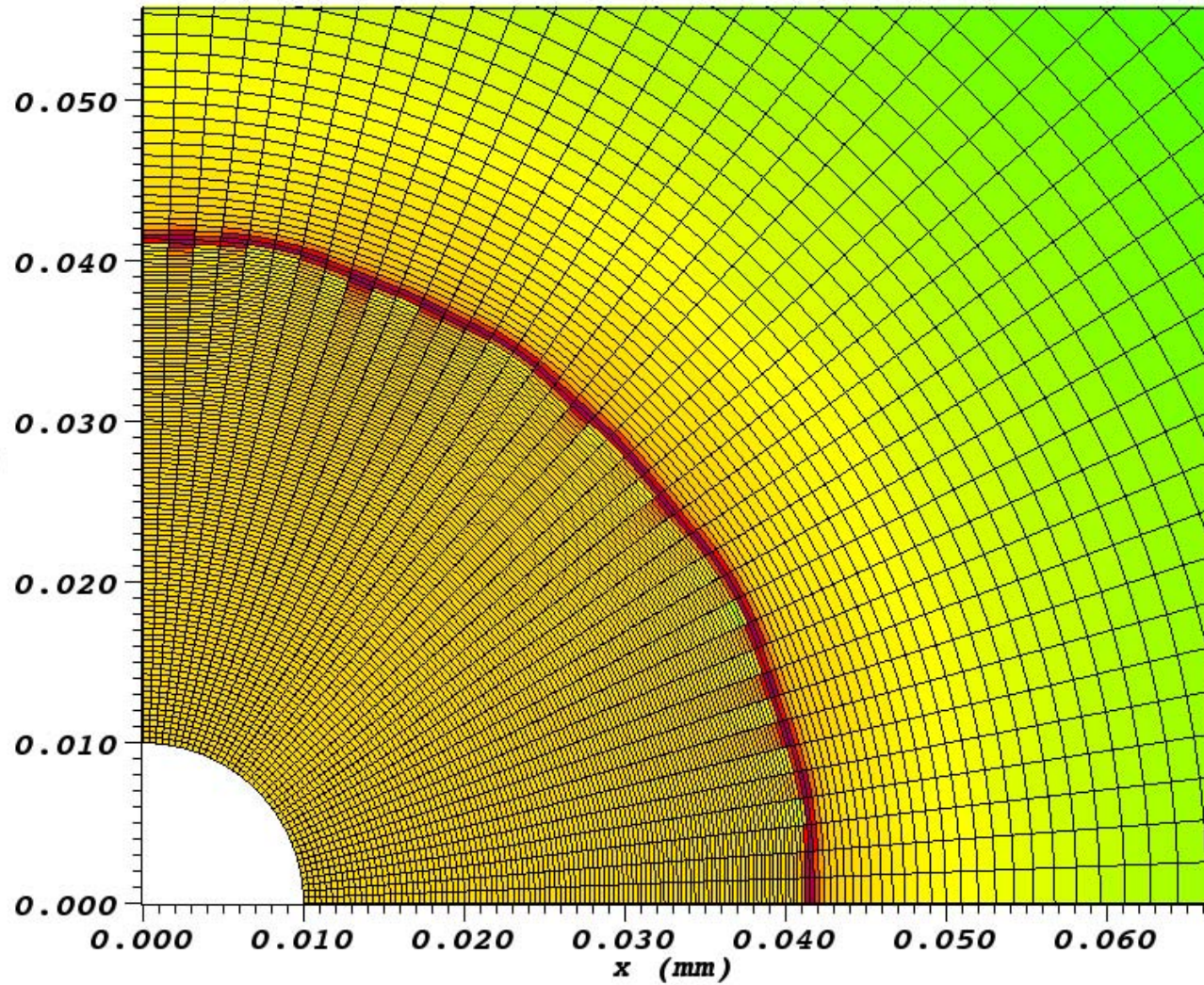
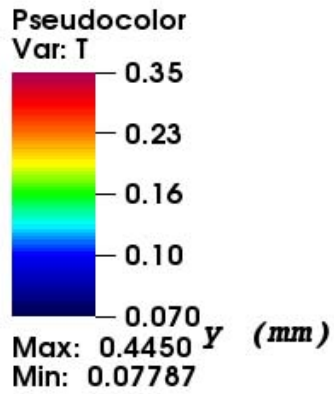
DB: vtkall010.vtk
Cycle: 10



user: Mikhail Basko
Sat Nov 07 12:15:10 2009

t = 3 ns

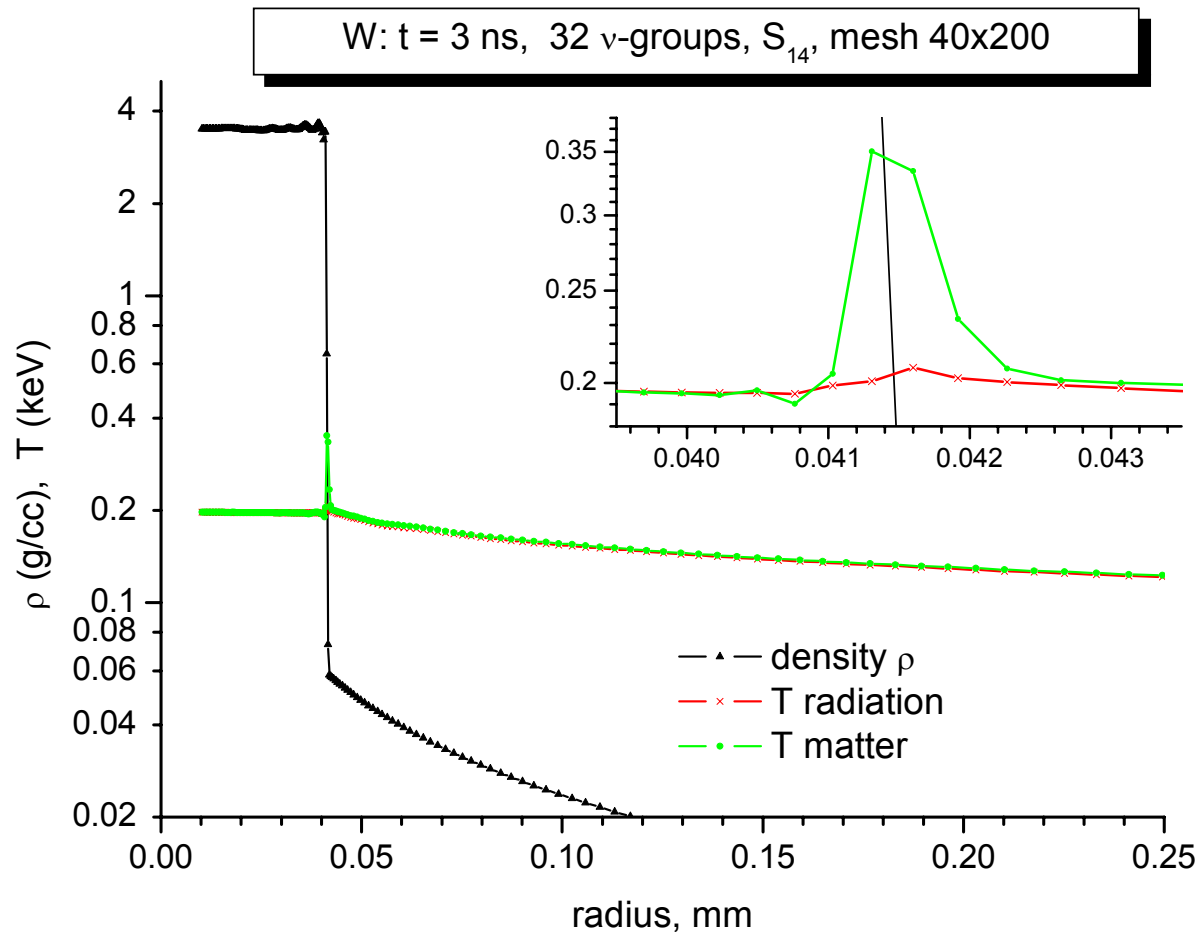
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Cycle: 6
Mesh
Var: mesh



Shock structure



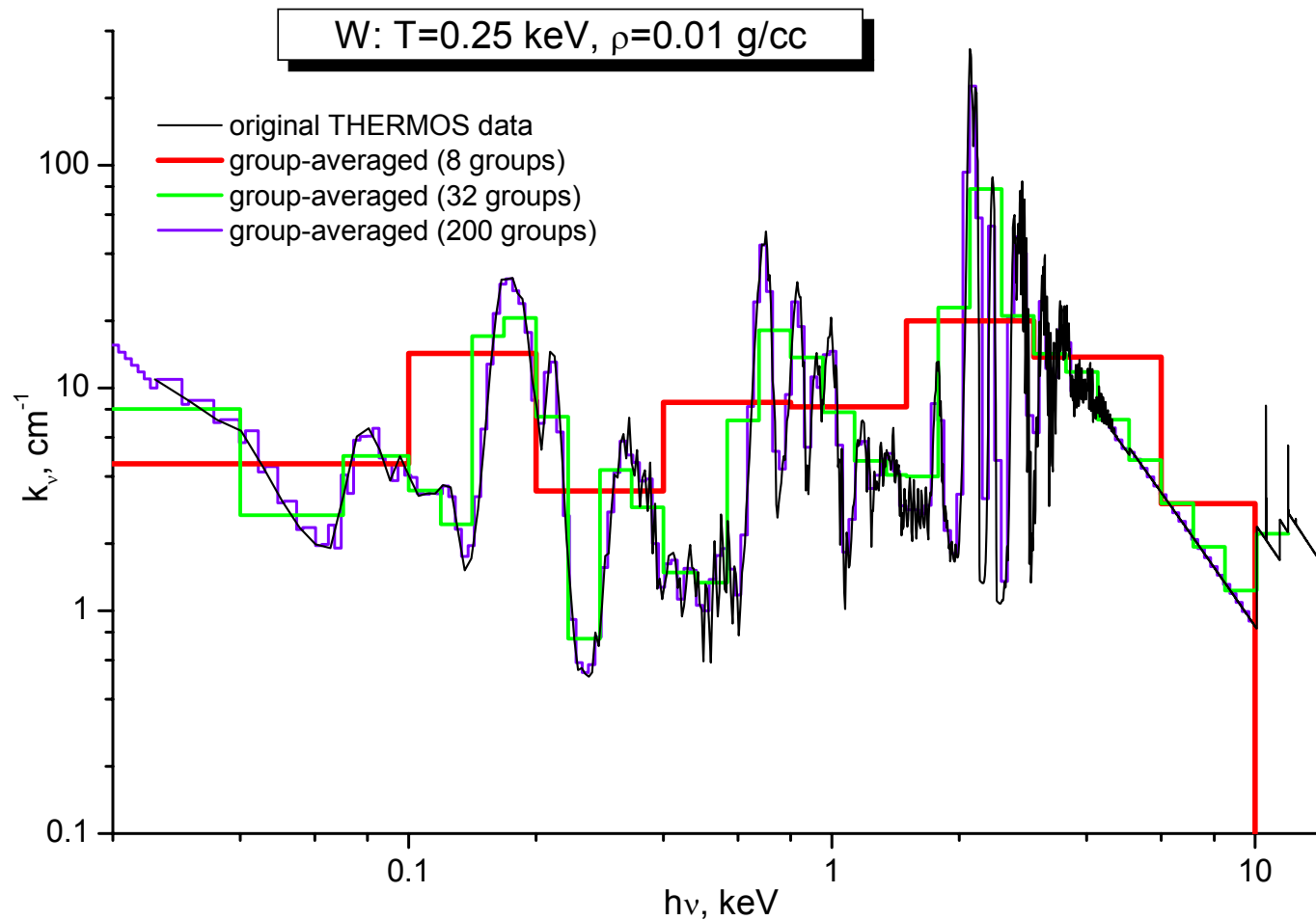
Here we deal with a radiation-dominated shock front, which has a supercritical amplitude (see Zel'dovich and Raizer, chapter VII). The shocked material radiates away about 90% of its initial kinetic energy.



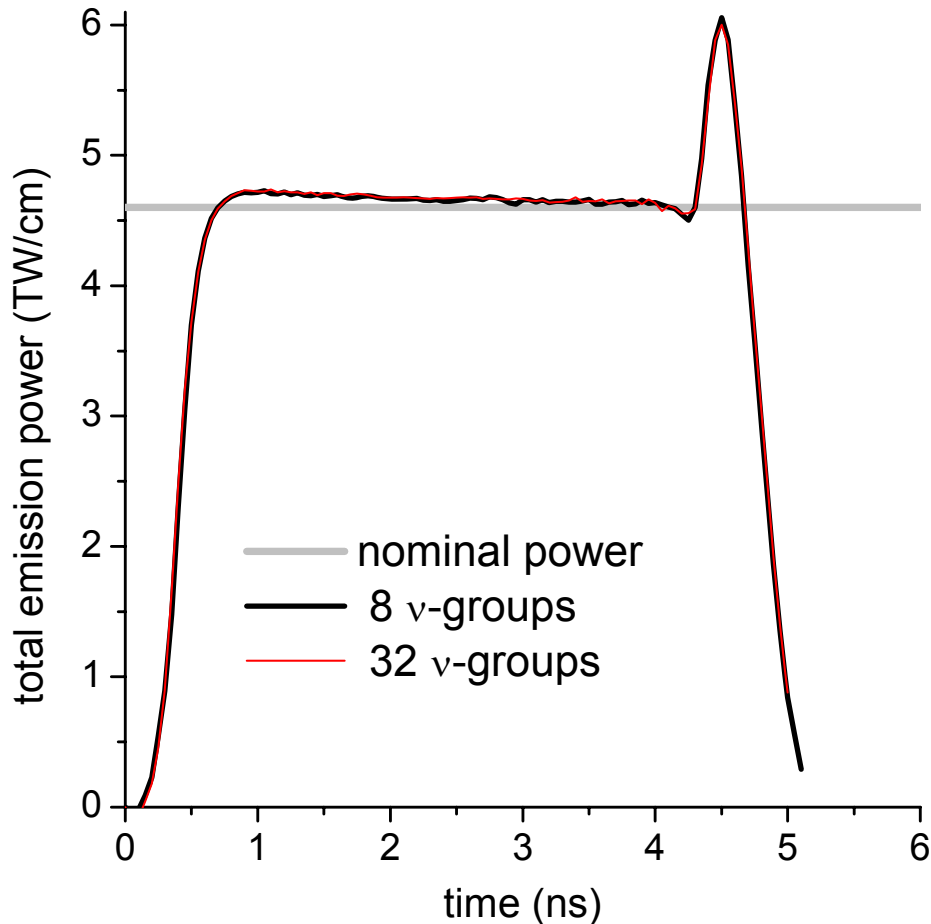
Plasma opacities



Plasma opacities for the above simulation have been provided by the group of V.G.Novikov from the Keldysh institute (KIAM) with the THERMOS code.



Total X-ray emission power



The imploding tungsten plasma radiates away about 90% of its initial kinetic energy.

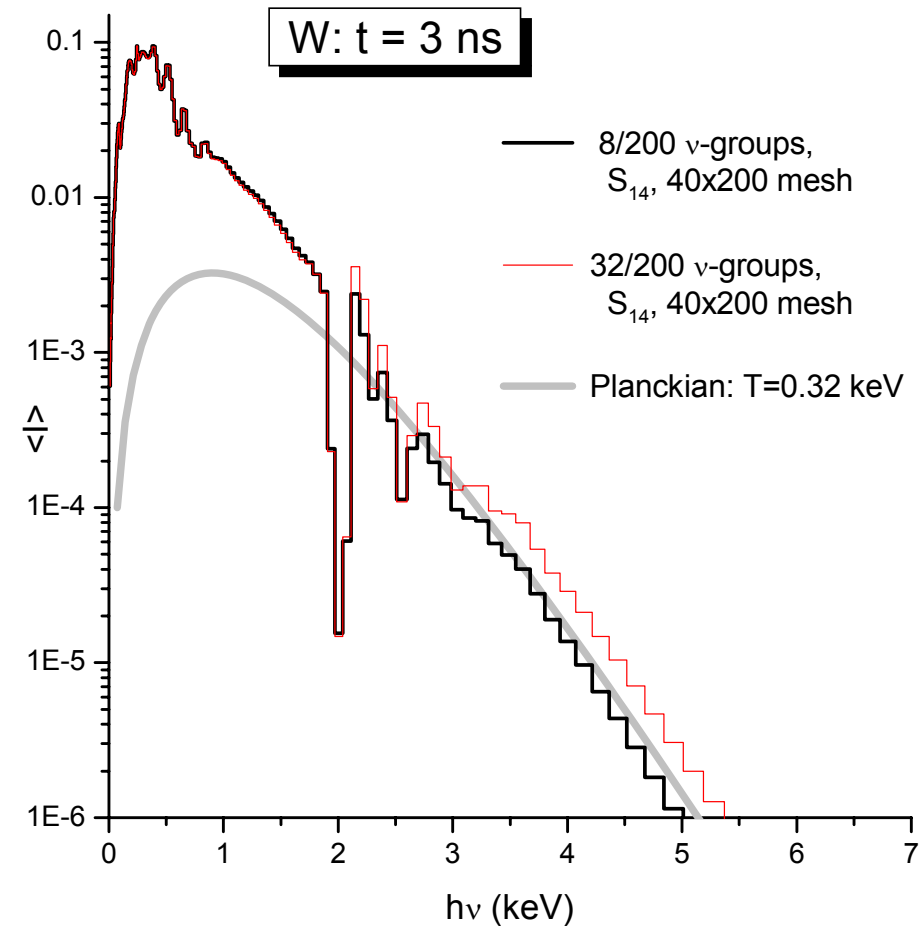
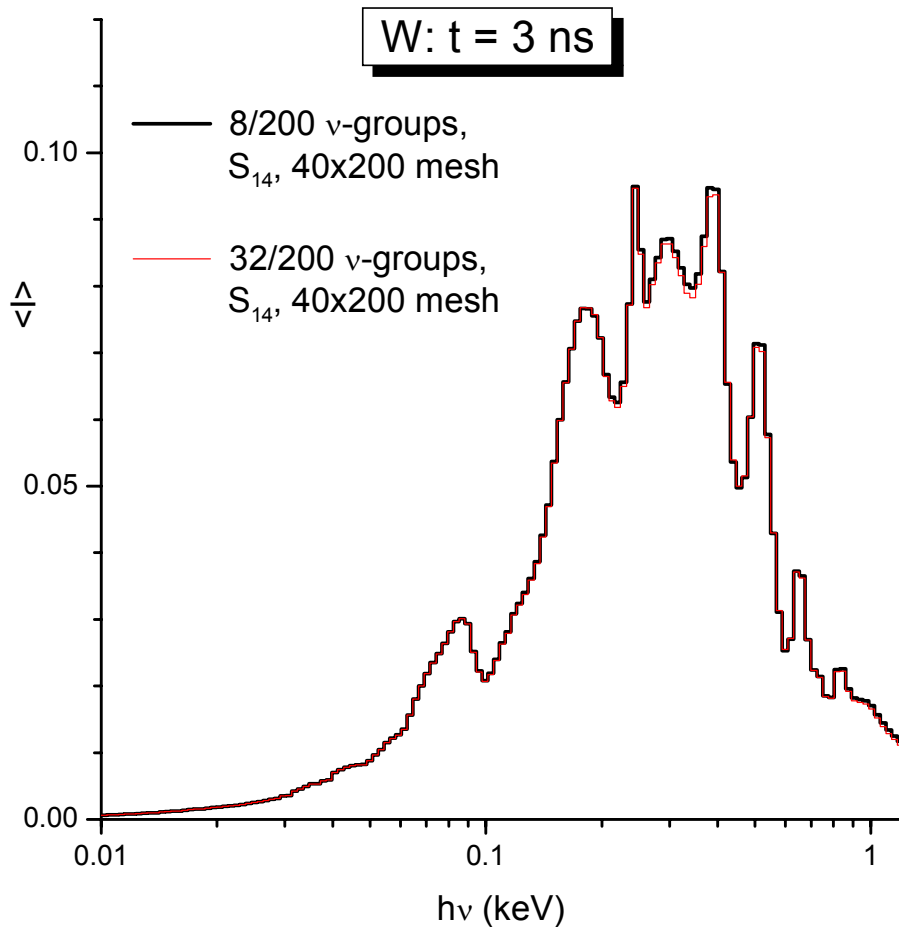
The nominal power of implosion is

$$\begin{aligned} W_{nom} &= \frac{1}{2} M U_{im}^2 / t_{pulse} = \\ &= 23 \text{ kJ cm}^{-1} / 5 \text{ ns} = 4.6 \text{ TW cm}^{-1} \end{aligned}$$

Spatially integrated emission spectra



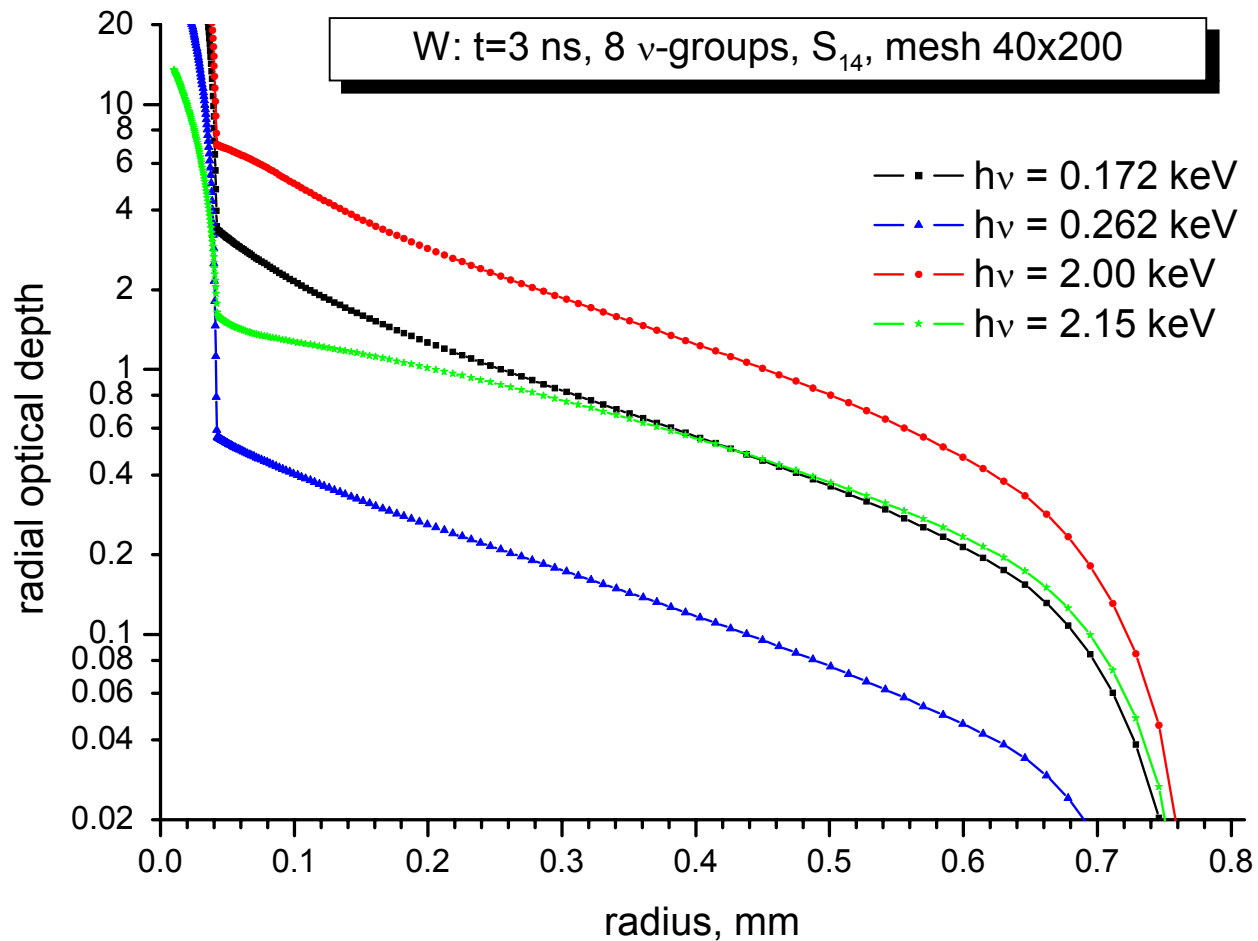
High-resolution emission spectra can be obtained in the post-processor regime even when hydrodynamics is simulated with a relatively small number (~ 10) of spectral groups.



Optical thickness of the imploding plasma



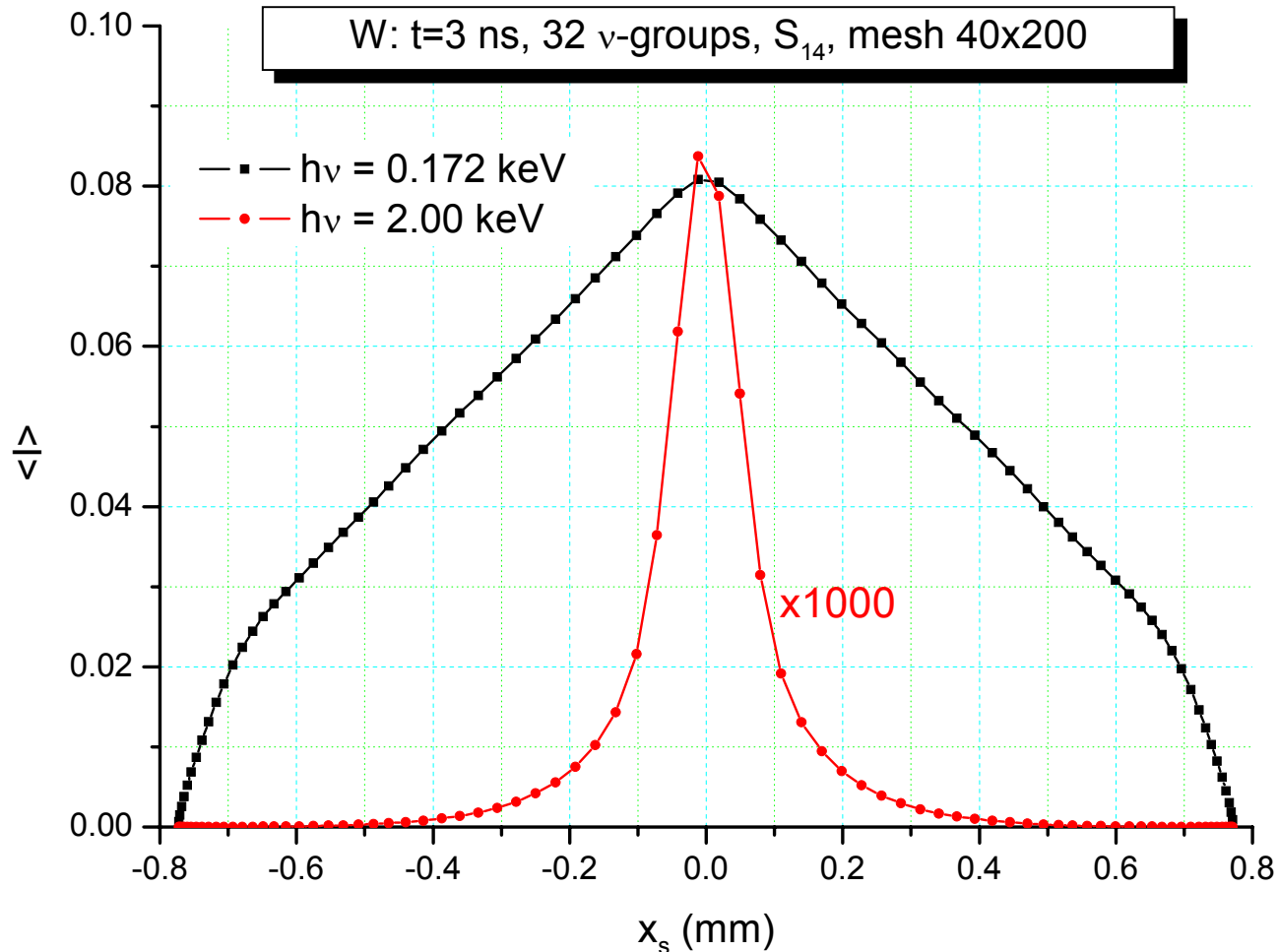
Over a wide range of frequencies, the shock front lies at an optical depth of $\tau = 0.5-2.0$
⇒ the multi-frequency diffusion is not an adequate approach for such situations!



Calculated X-ray images of the central pinch



Spectral X-ray images provide information on the internal structure of the imploding pinch.



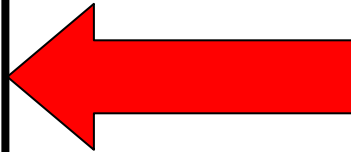
**Problem 2: Hole boring in a copper foil by an
intense nanosecond laser pulse**

Copper foil under PHELIX (GSI) irradiation

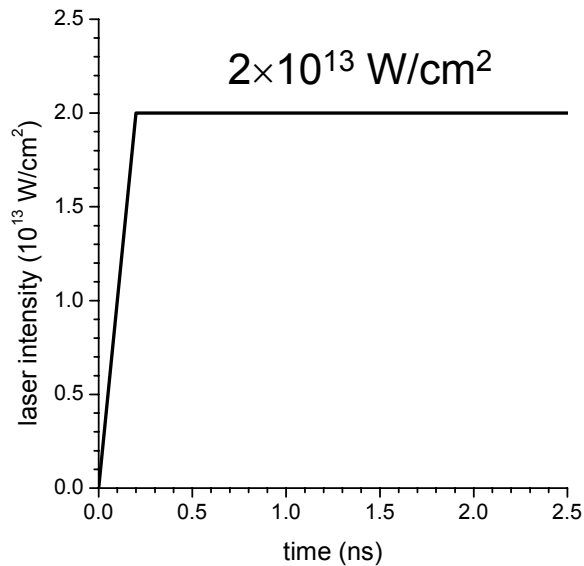


Cu foil: $4\text{g/cc} \times 2\mu\text{m} = 0.8\text{ mg/cm}^2$, $T_0 = 0.58\text{ eV}$

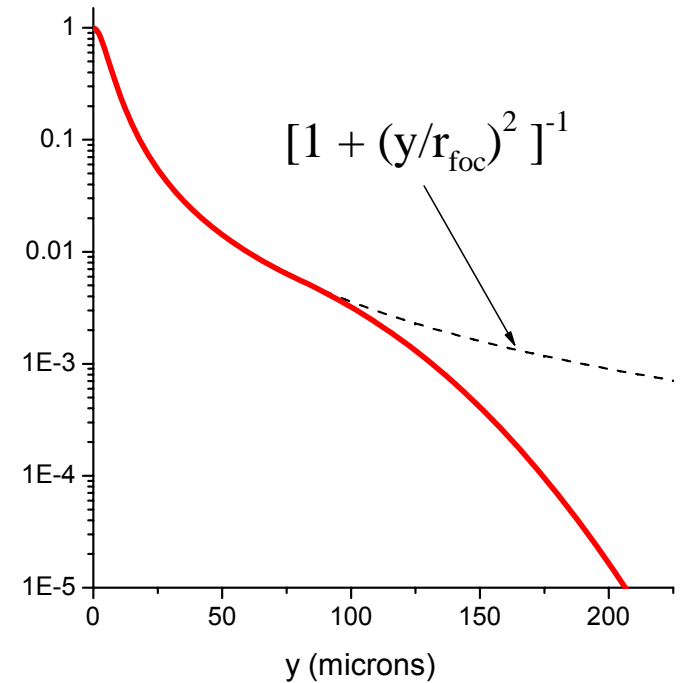
PHELIX beam (prepulse): 2-3 ns with $2 \times 10^{13}\text{ W/cm}^2$



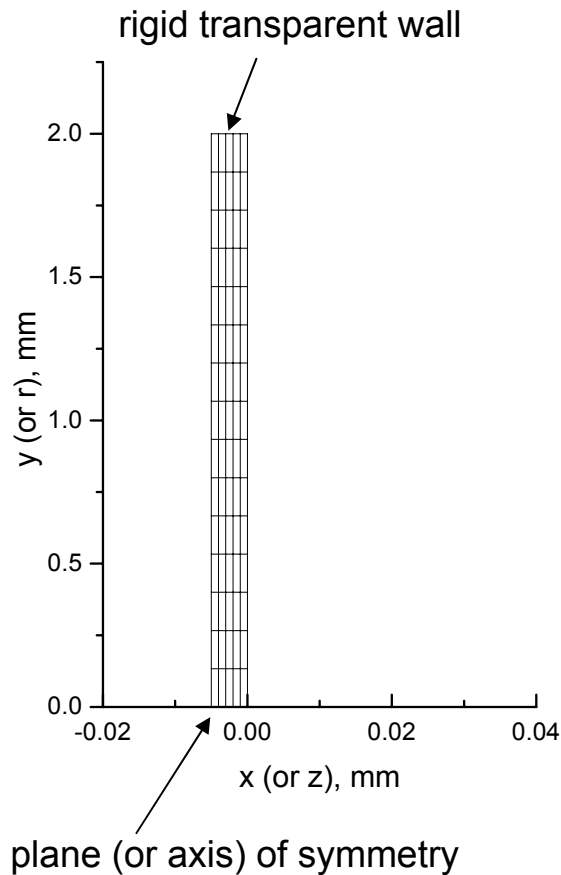
Temporal pulse profile



Spatial pulse profile



Simulation setup



Initial conditions: $\rho_0 = 4.0 \text{ g/cc}$, $\Delta x = 2 \text{ }\mu\text{m}$,
 $p_0 = 3 \text{ kbar}$, $T_0 = 0.58 \text{ eV}$

Boundary conditions: $p_b = 3 \text{ kbar}$

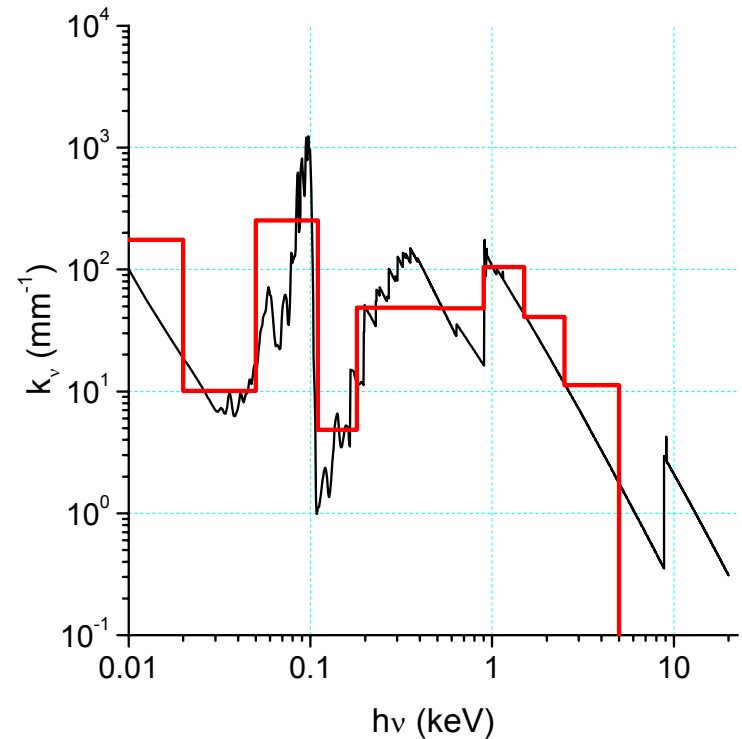
EOS: Novikov *et al.*

Opacity: Novikov *et al.*
9 freq. groups

Laser deposition: transport
of one beam with the inverse
bremsstrahlung absorption
coefficient

Mesh: $n_x \times n_y = 150 \times 300$

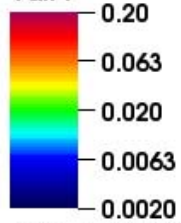
Spectral opacity of Cu



DB: vtkall0000.vtk
Cycle: 0

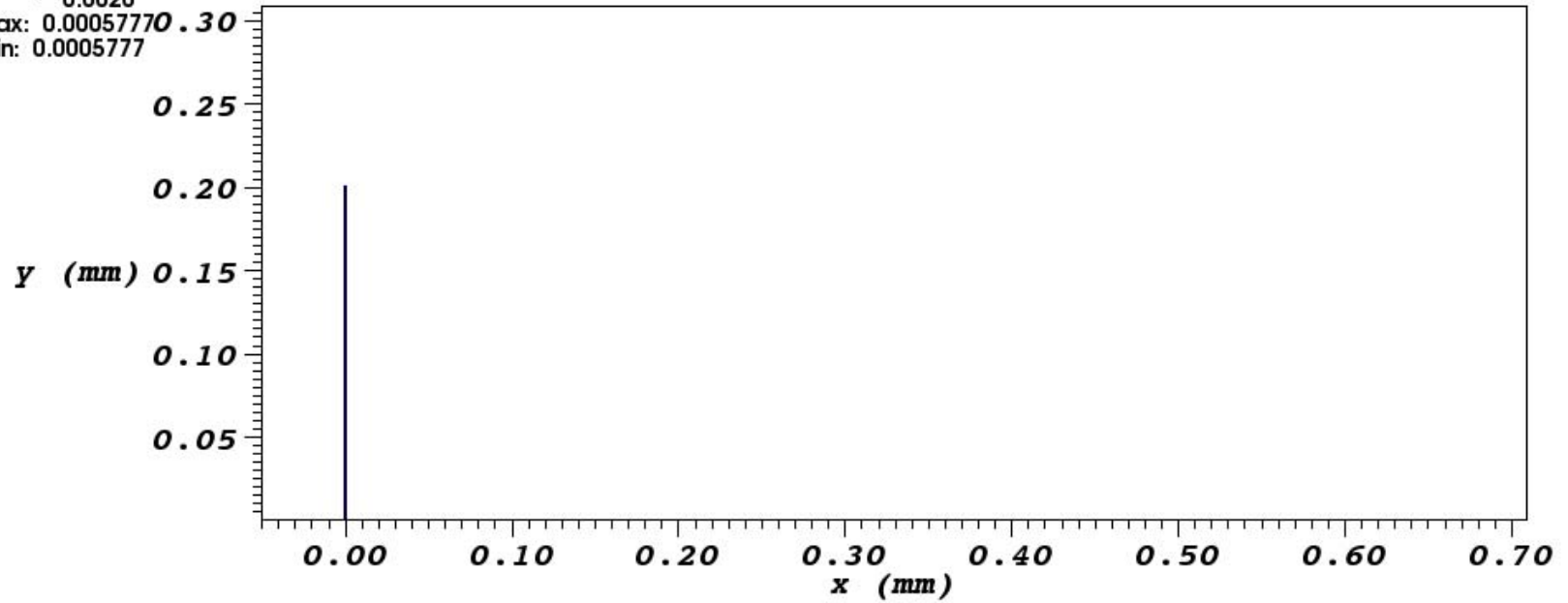
Pseudocolor

Var: T



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Min: 0.0005777

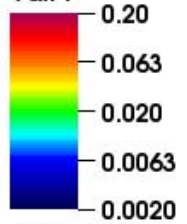
Temperature



user: Mikhail Basko
Wed Nov 25 14:50:16 2009

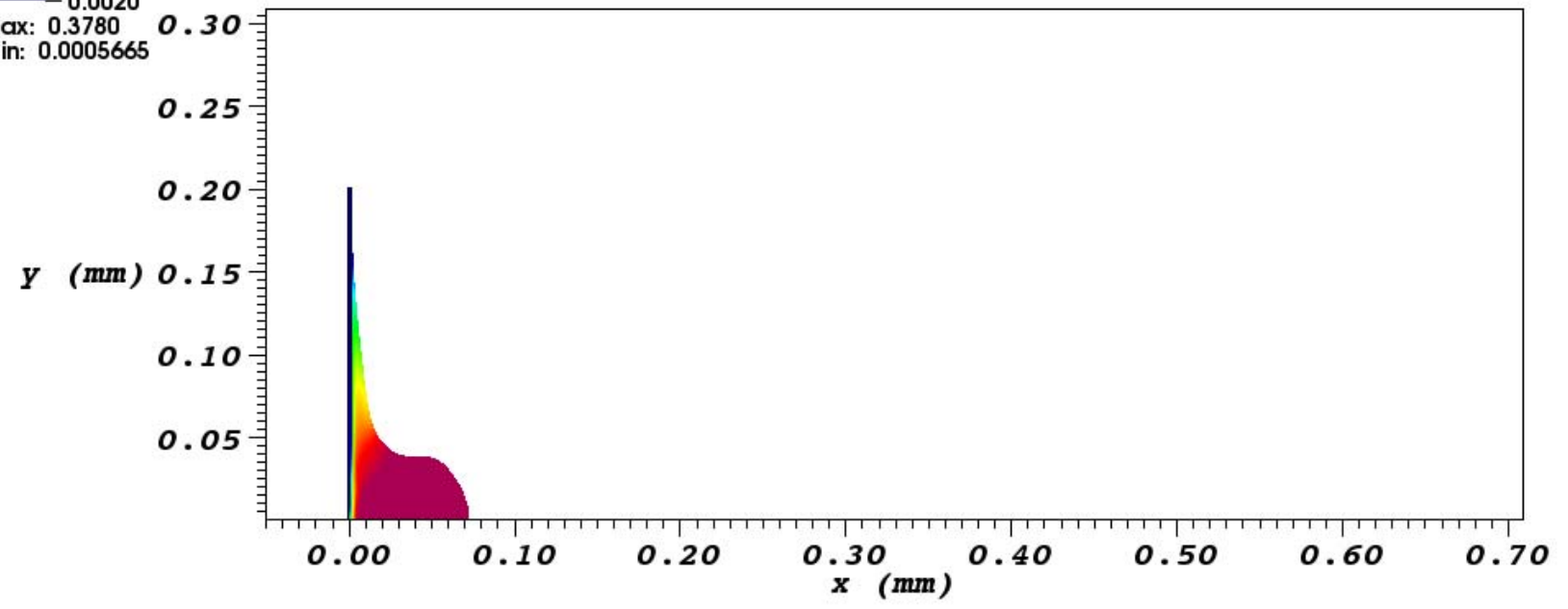
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Cycle: 20

Pseudocolor
Var: T



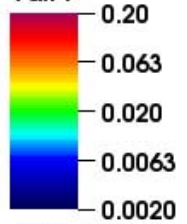
Max: 0.3780
Min: 0.0005665

Temperature



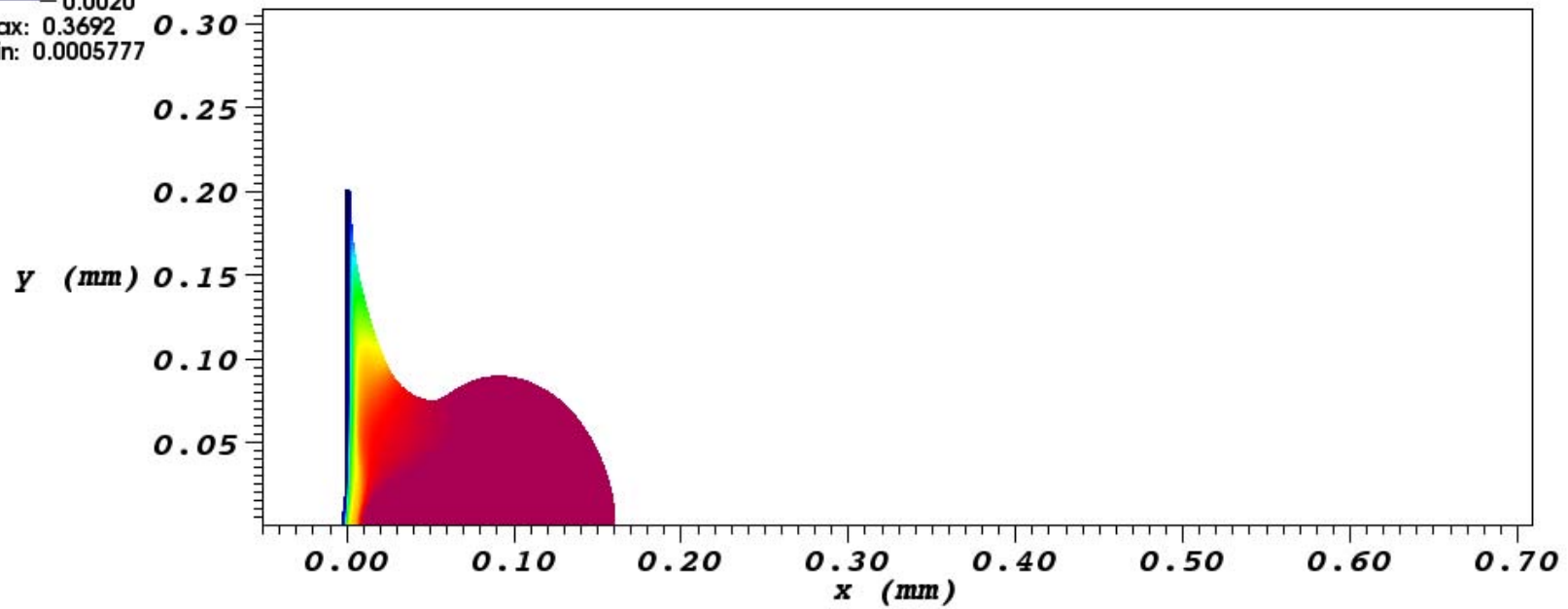
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Cycle: 40

Pseudocolor
Var: T



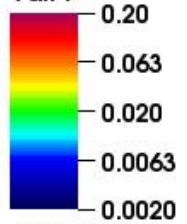
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Min: 0.0005777

Temperature



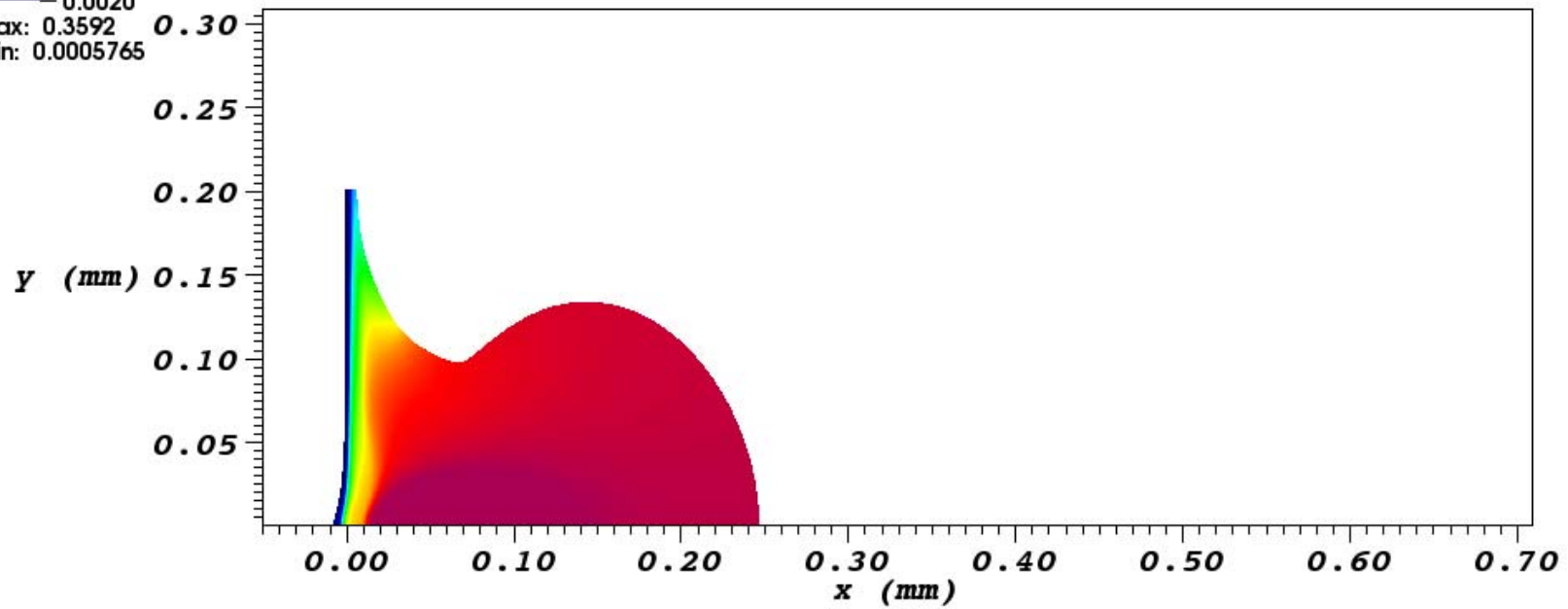
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Cycle: 60

Pseudocolor
Var: T



Max: 0.3592
Min: 0.0005765

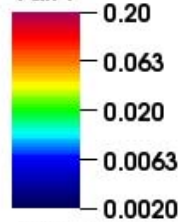
Temperature



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Cycle: 80

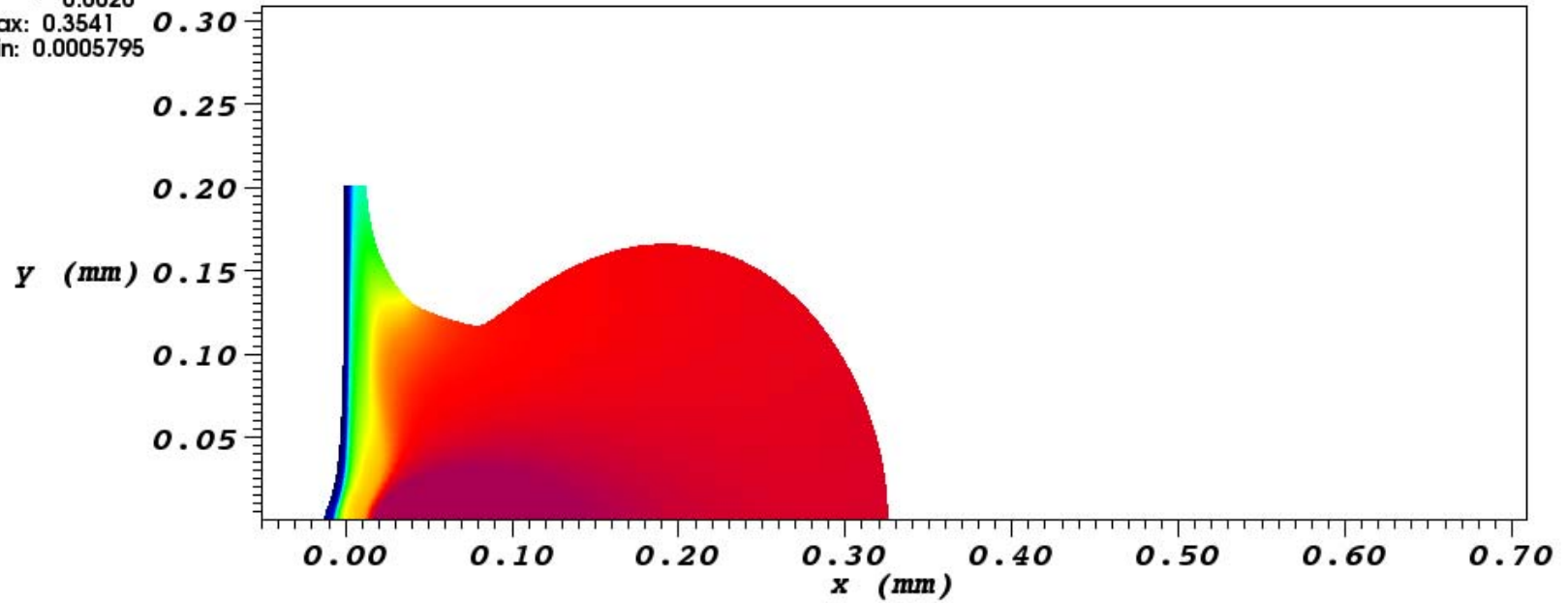
Pseudocolor

Var: T



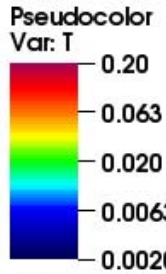
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Min: 0.0005795

Temperature



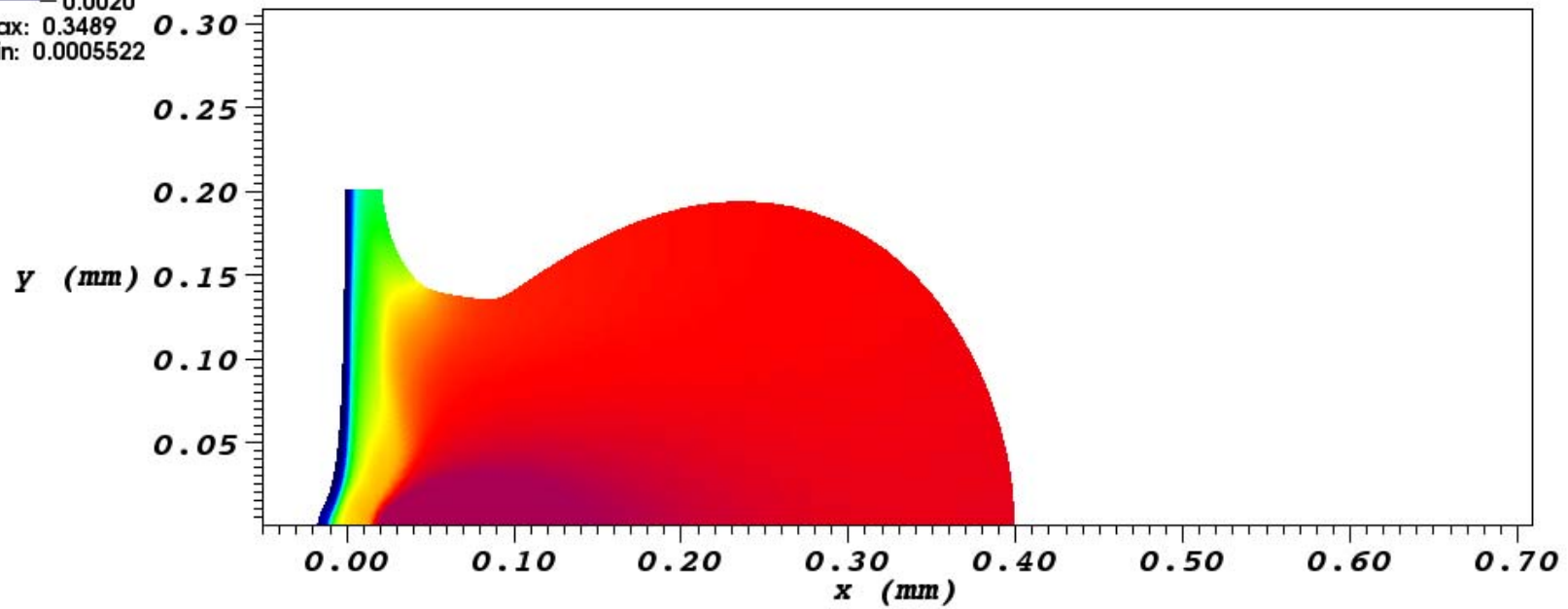
user: Mikhail Basko
Wed Nov 25 14:52:18 2009

DB: vtkall0100.vtk
Cycle: 100



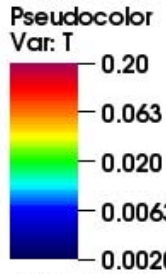
Max: 0.3489
Min: 0.0005522

Temperature



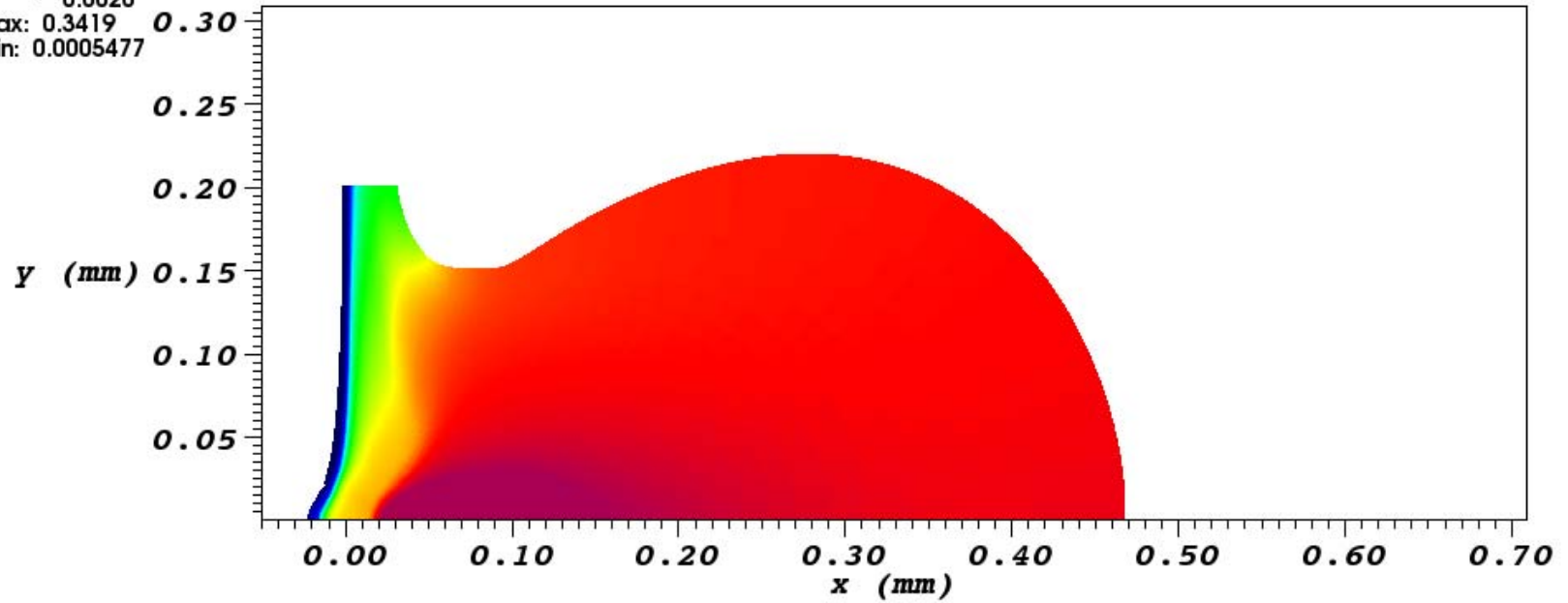
user: Mikhail Basko
Wed Nov 25 14:52:27 2009

DB: vtkall0120.vtk
Cycle: 120



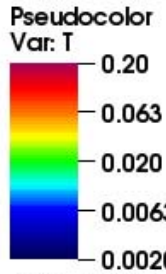
Max: 0.3419
Min: 0.0005477

Temperature



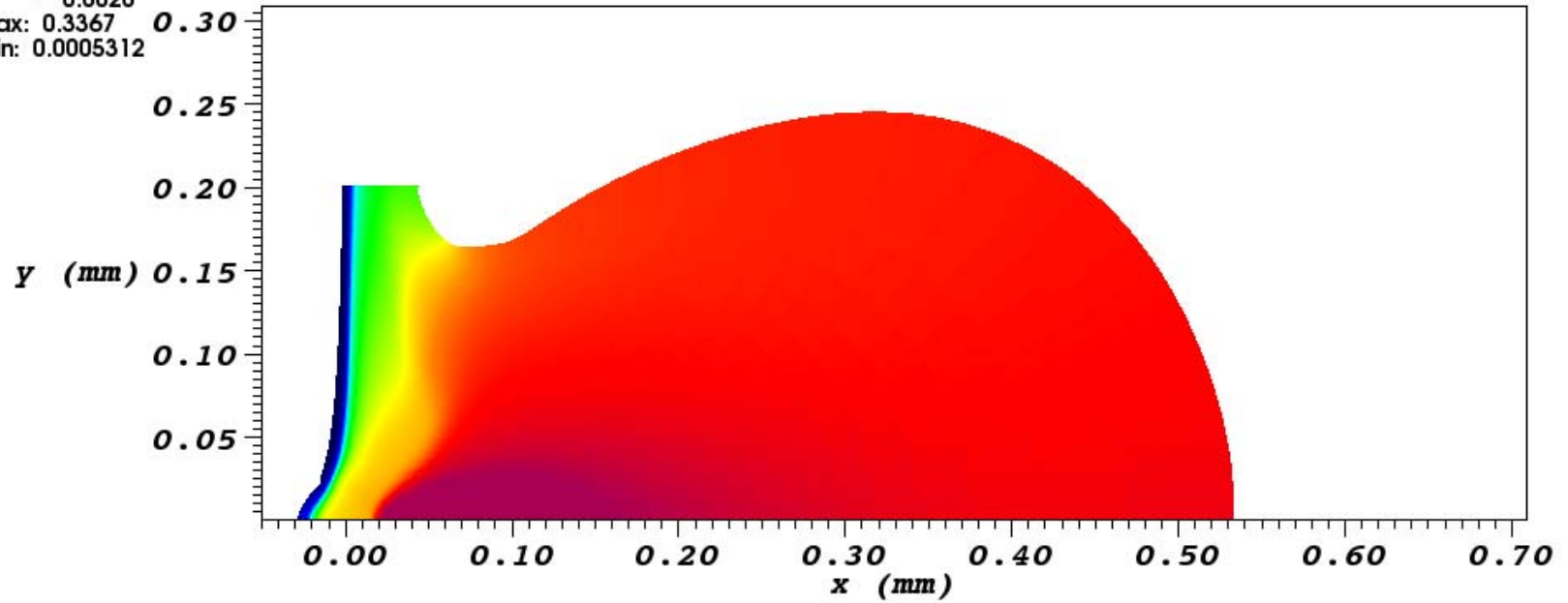
user: Mikhail Basko
Wed Nov 25 14:52:38 2009

DB: vtkall0140.vtk
Cycle: 140



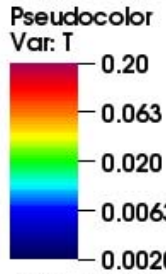
Max: 0.3367
Min: 0.0005312

Temperature



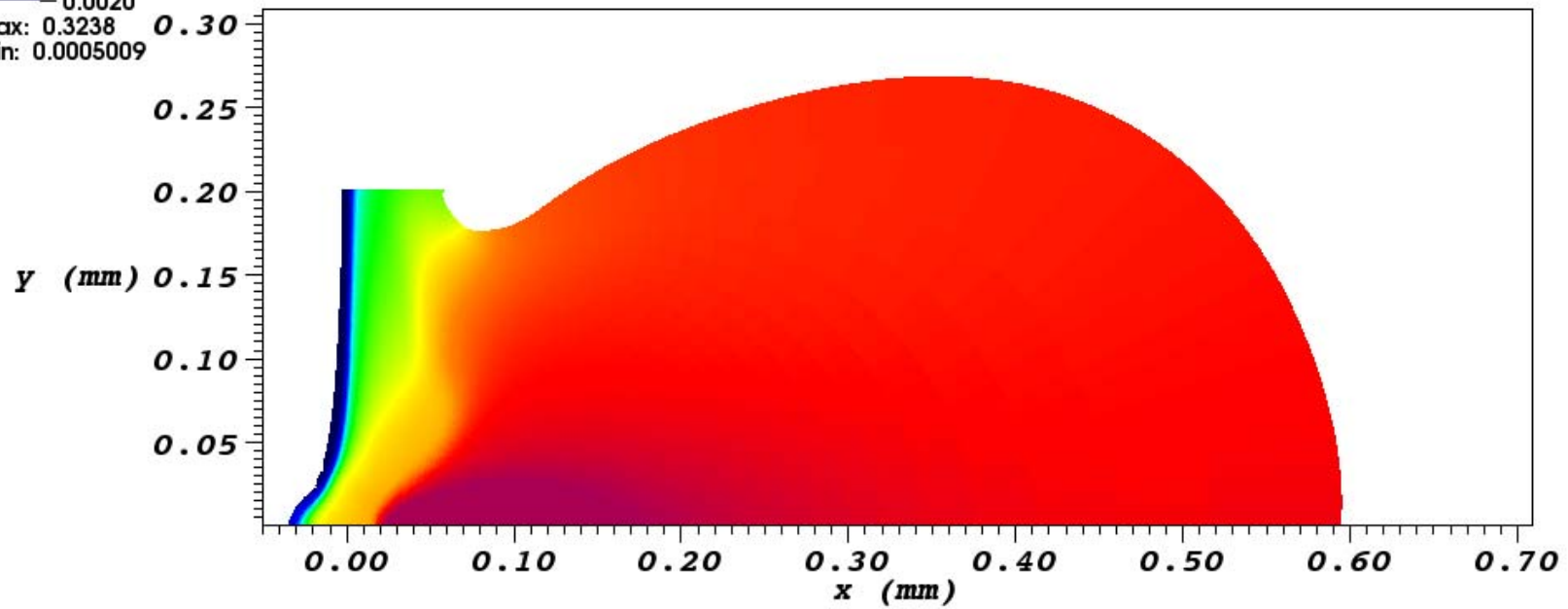
user: Mikhail Basko
Wed Nov 25 14:52:50 2009

DB: vtkall0160.vtk
Cycle: 160



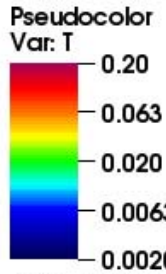
Max: 0.3238
Min: 0.0005009

Temperature



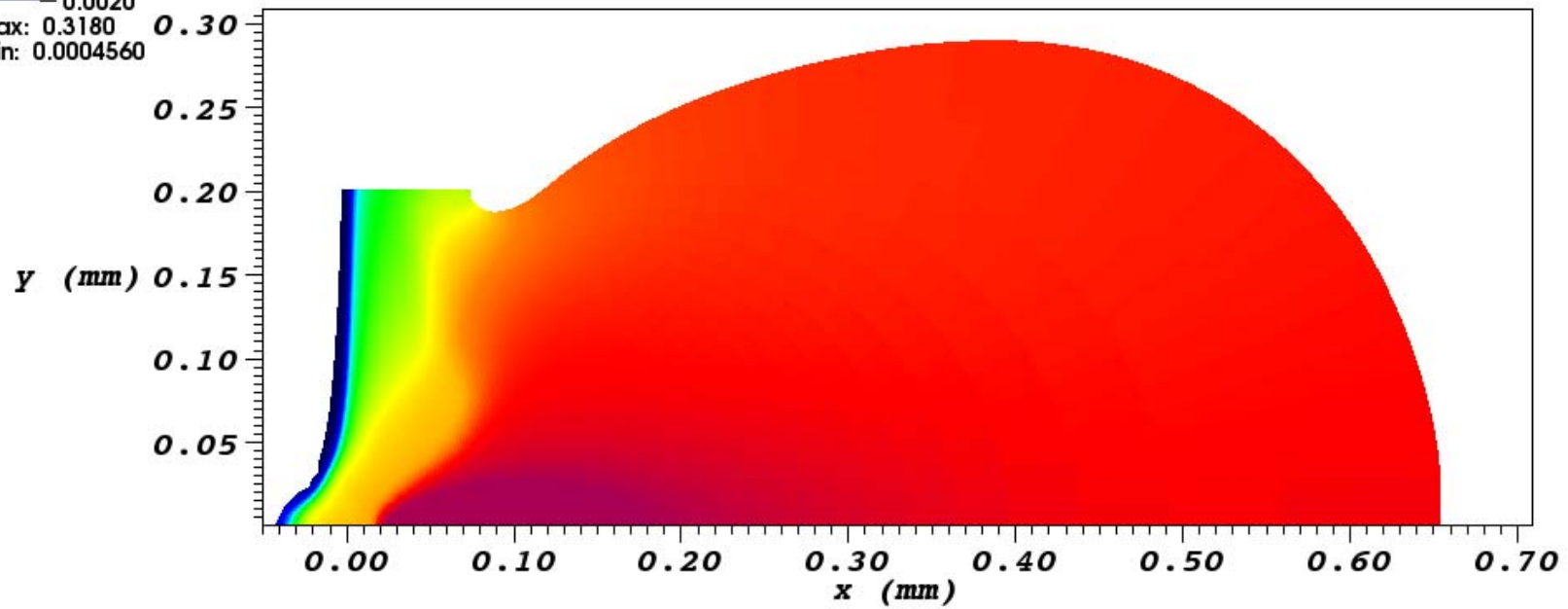
user: Mikhail Basko
Wed Nov 25 14:53:01 2009

DB: vtkall0180.vtk
Cycle: 180



Max: 0.3180
Min: 0.0004560

Temperature

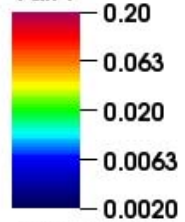


user: Mikhail Basko
Wed Nov 25 14:53:11 2009

DB: vtkall0200.vtk
Cycle: 200

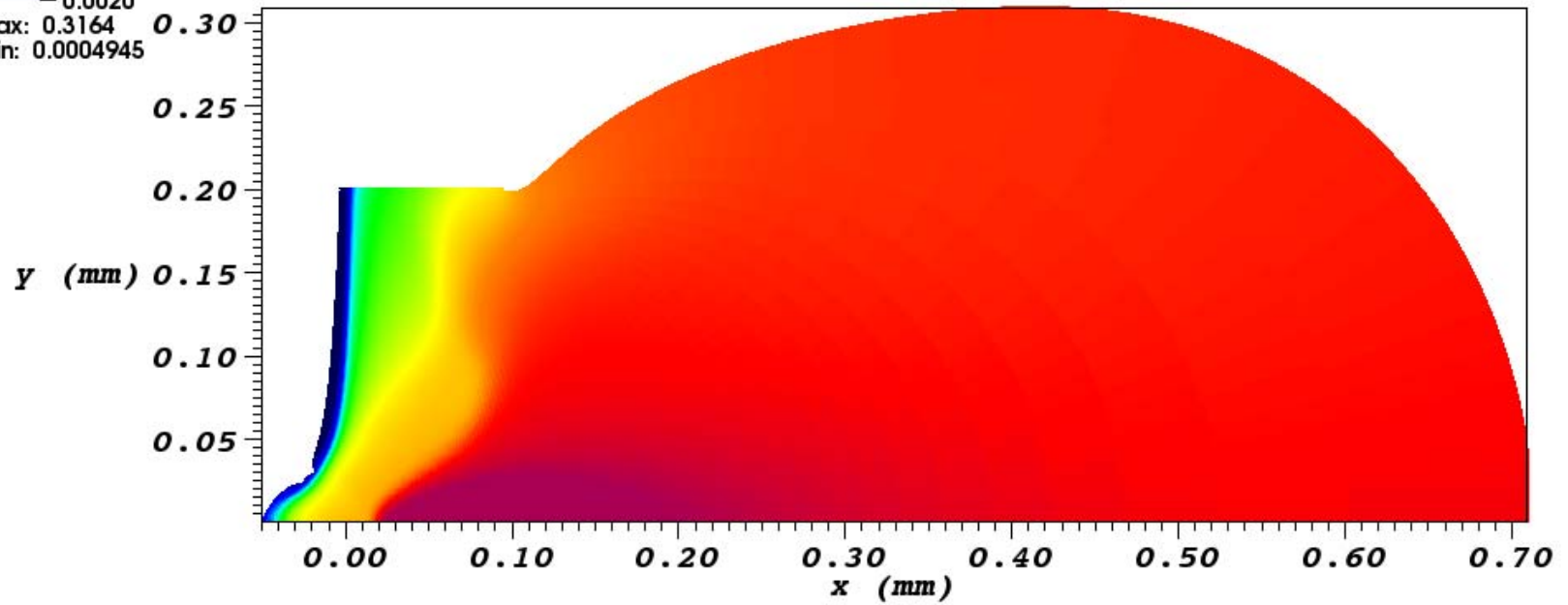
Pseudocolor

Var: T

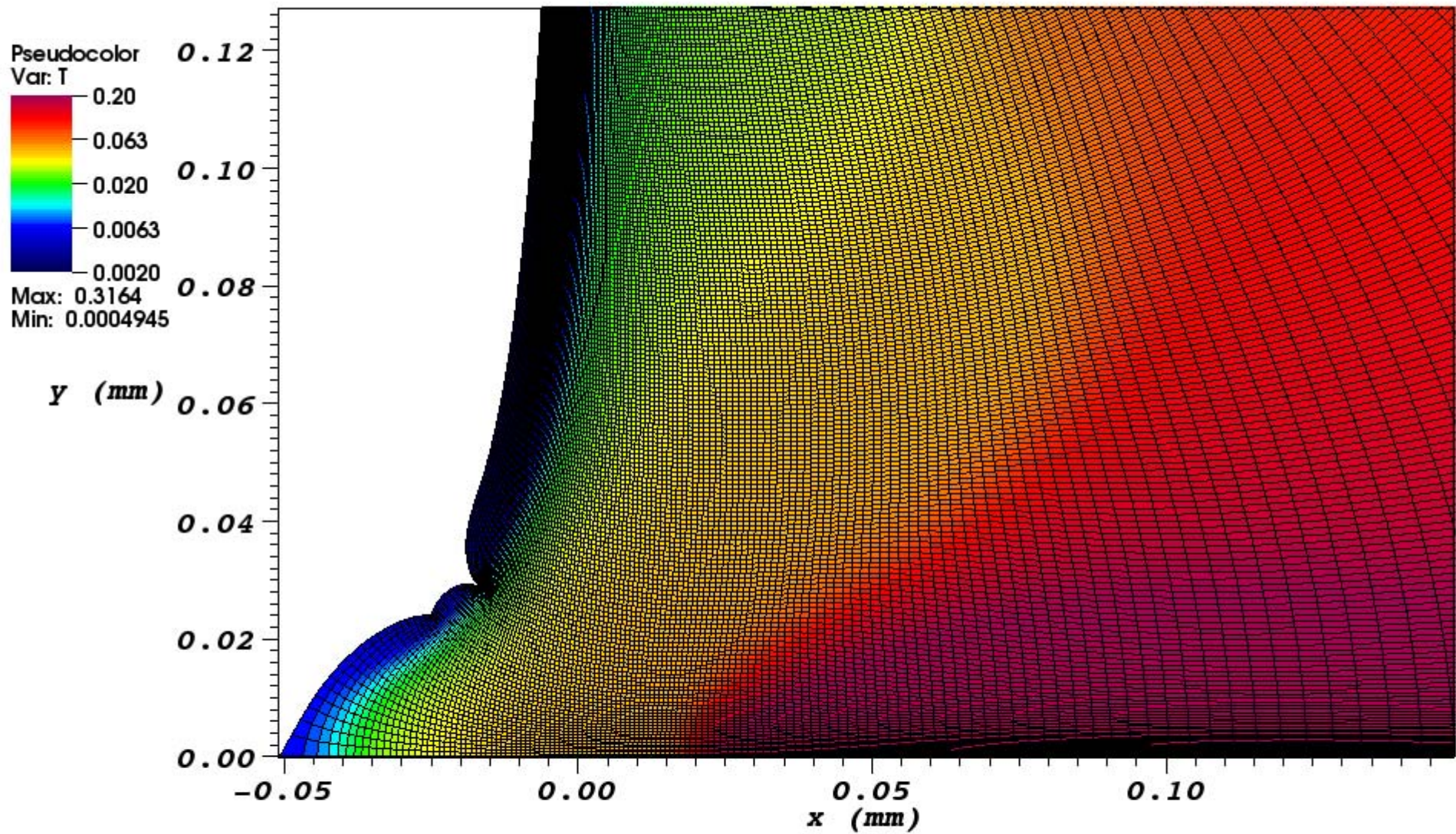


Max: 0.3164
Min: 0.0004945

Temperature



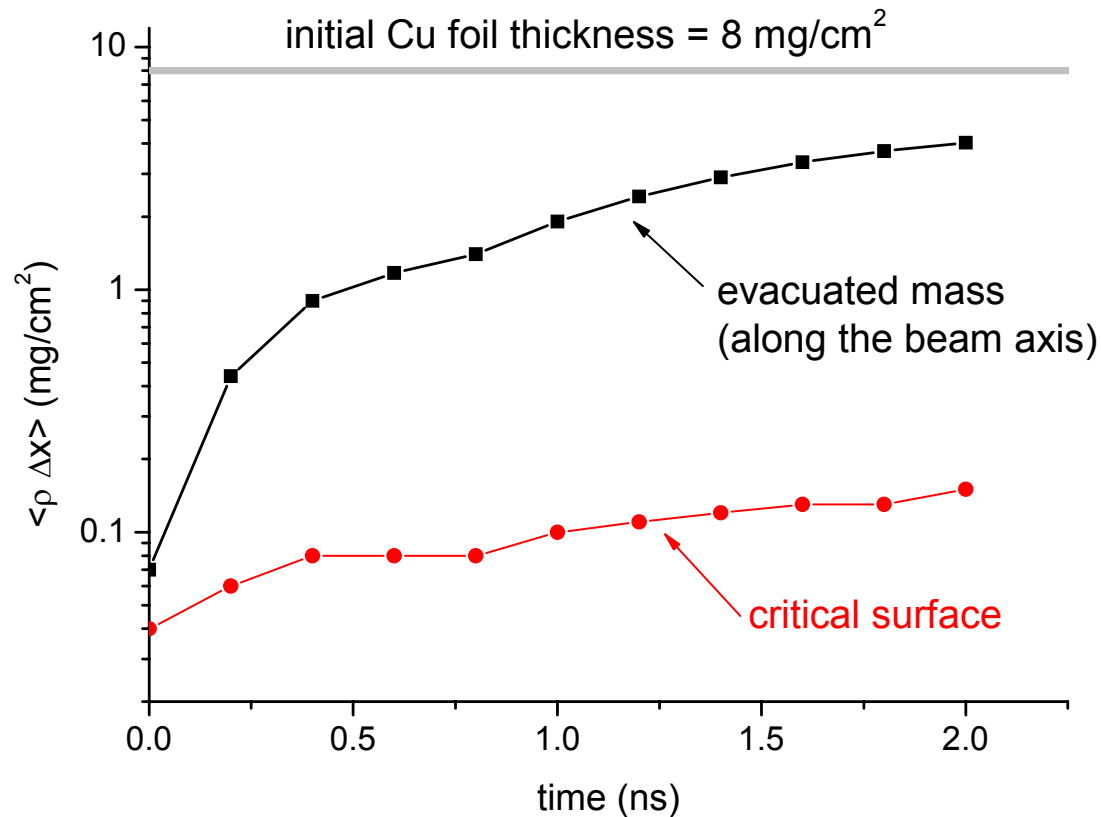
user: Mikhail Basko
Wed Nov 25 14:53:28 2009



Effects of 2D expansion



Time dependence of mass column density $\int \rho dx$ along the central axis at $y = 0$.



The effect of hole boring:
the evacuated mass exceeds the depth of the critical surface by as large a factor as 20–30.

Conclusions



The status of the RALEF-2D code:

- ❖ the first results are encouraging,
- ❖ the work on further development continues.