M.I. Polikarpov, ITEP, Moscow

Introduction: quarks, gluons and QCD Millennium problem: Confinement Supercomputers and strong interactions Heavy ion collisions and computers Graphene and computers

Physics of high energy density in matter

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Physics of high energy density in matter

Three examples of computer simulations of strongly interacting systems

1. Confinement problem in QCD



2. Interference of strong and electromagnetic interactions in heavy ion collisions



3. Graphene as quantum field theory



Experiment LHC RHIC







 \mathbf{C}_{2}

Σ



Supercalculations



Interactions – 1. Gravity



mg

Interactions – 2. Weak



Interactions – 3. Electromagnetism





Interactions – 4. Strong



Main problems of strong interaction theory, QCD

Derive from QCD Lagrangian

$$L = -\frac{1}{g^2} \operatorname{Tr} F_{\mu\nu}^2 + \sum_{f} \bar{\psi}_{f} (D+m) \psi_{f}$$

- (1) Hadron spectrum,(2) Matrix elements,(3) Phase diagram
- (4) Explain color confinement



http://www.claymath.org/millennium/Yang-Mills Theory/ (1 000 000 \$US)

Color confinement

(Why we do not observe free quarks and gluons?)

The main difficulty is the absence of first-principle nonperturbative methods in QCD. Computers can prove confinement "numerically"



Force between quark and antiquark is12 tons!!!

http://www.claymath.org/millennium

Quantum mechanics of a particle



Quantum field theory

 $A_{\mu}(x) = A_{\mu}(x, y, z, t)$

 $-\infty < A_{\mu}(x) < +\infty$

 $Z = \iiint DA_{\mu}(x) e^{iS[A_{\mu}]}$

Methods

• Imaginary time *t*→*it*

X

$$Z = \int D\varphi \exp\{i S[\varphi]\} \longrightarrow Z = \int D\varphi \exp\{-S[\varphi]\}$$

• Space-time discretization
$$D\varphi(x) \Rightarrow \prod d\varphi_x \qquad \qquad Z = \int \prod d\varphi_x \exp\{-S[\varphi]\}$$

• Thus we get from functional integral the partition function for statistical theory in four dimensions

X

INTRODUCTION Three limits



Lattice spacing

Lattice size

Quark mass



Typical values

 $a \approx 0.1 \ fm$ $L \approx 2 \div 4 \ fm$ $m_q \approx 100 \ Mev$



Typical multiplicity of integrals

For lattice L⁴ (*L*=48, *L*⁴=5,308,416)

Te multiplicity of integralsover gluon fields is 32L⁴ (L=48, 32L⁴=169,869,312)

• For quark fields we work with matrices 12L⁴ x 12L⁴ (L=48, 12L⁴=63,700,992)

$$\int d\psi \, d\overline{\psi} \, \exp\{\overline{\psi}M\psi\} = \det M$$

SU(2) glue SU(3) glu

The force between quark and antiquark is 12 tons!!!



SU(2) glue SU



SU(2) glue SU(3) glue 2qQCD (2+1)QCD Three body forces!





$$V(r_1, r_2, r_3) \neq V(r_1 - r_2) + V(r_2 - r_3) + V(r_3 - r_1)$$



The origine of the mass











<u>Masses of material objects is due to gluon</u> <u>fields inside baryon</u>



 $E = m_0 c^2$

 $3m_q / m_{baryon} \approx 1/100$



Three body forces!





Figure 9: The monopole part of the baryon potential at finite temperature in full QCD as a function of $L_{\mathbf{Y}}$ ($T < T_c$) and L_{Δ} ($T > T_c$), respectively, in units of God knows what.

In Fig. 9 we show the baryon potential on the $16^3 8$ lattice at $\beta = 5.2$ for several values of κ . At this β value

$$T \propto \exp(-2.81/\kappa)$$
. (4.1)

Increasing κ thus increases the temperature. We cross the finite temperature phase transition at $\kappa = 0.1344$ [14]. We see that the potential flattens off while we approach the transition point. However, the distances we were able to probe are not large enough to make any statement about string breaking.

To compute the action density ρ_A^{3Q} and the electric field and monopole correlators E_i^{3Q} and k^{3Q} , respectively, we need to reduce the statistical noise. Note that the Polyakov loops span an area of $\approx 16 \times 8$ lattice spacings. We do that by using extended operators

$$\begin{split} {}^{3Q}_{A}(s) &\longrightarrow \frac{1}{8} \{ \rho_{A}^{3Q}(s) + \rho_{A}^{3Q}(s - \hat{x} - \hat{y} - \hat{z}) + \rho_{A}^{3Q}(s - \hat{x} - \hat{y}) \\ &+ \rho_{A}^{3Q}(s - \hat{x} - \hat{z}) + \rho_{A}^{3Q}(s - \hat{y} - \hat{z}) + \rho_{A}^{3Q}(s - \hat{x}) \\ &+ \rho_{A}^{3Q}(s - \hat{y}) + \rho_{A}^{3Q}(s - \hat{z}) \} \,. \end{split}$$

$$\end{split}$$

$$(4.2)$$

$$E_i^{3Q}(s) \longrightarrow \frac{1}{4} \{ E_i^{3Q}(s) + E_i^{3Q}(s - \hat{x} - \hat{t}) \\ + E_i^{3Q}(s - \hat{x}) + E_i^{3Q}(s - \hat{t}) \},$$
(4.3)

$$k^{3Q}(*s,\mu) \longrightarrow \frac{1}{2} \{ k^{3Q}(*s,\mu) + k^{3Q}(*(s-\hat{z}),\mu) \}, \qquad (4.4)$$

where (again) we have assumed that the quarks lie in the (x, y) plane, and we call the direction of the Polyakov lines the t direction.

Usually the teams are rather big, 5 - 10 - 15 people

arXiv:hep-lat/0401026v1

arXiv:hep-lat/0401026v2



Figure 9: The monopole part of the baryon potential at finite temperature in full QCD as a function of L_Y ($T < T_c$) and L_Δ ($T > T_c$), respectively in units of God knows what.

In Fig. 9 we show the baryon potential on the $16^3 8$ lattice at $\beta = 5.2$ for several values of κ . At this β value

SU(2) glue SU(3) glue 2qQCD (2+1)QCD String Breaking (DIK collaboration)





Hadron Mass Spectrum

angular momentum. For Λ and Σ resonances, the symbol is $L_{1,2J}$.

Meson Summary Table

Baryon Summary Table

See also the table of suggested qq quark-model assignments in the Quark Model section.

Indicates particles that appear in the preceding Meson Summary Table. We do not regard the other entries as being established.
 † Indicates that the value of J given is preferred, but needs confirmation.

	LIGHT UN	FLAVORED		STRA	NGE	BOTTOM		
	$(S = C \pm B = 0)$			$(S = \pm 1, C)$	= B = 0)	$(B = \pm 1)$		
	P ⁽ (<i>J</i> ^{, c})		1°(J' °)		1(3')		P(J'C)	
• π^{\pm}	1-(0-)	 π₂(1670) 	1-(2-+)	• K±	1/2(0-)	• B [±]	1/2(0 ⁻)	
• π ⁰	$1^{-}(0^{-+})$	 φ(1680) 	0-(1)	• K ⁰	1/2(0-)	• B ⁰	1/2(0)	
• <i>η</i>	0+(0 - +)	 ρ₃(1690) 	1+(3)	• K ⁰ _S	1/2(0-)	● <i>B</i> [±] / <i>B</i> ⁰ ADM	IXTURE	
 f₀(600) 	0+(0++)	 ρ(1700) 	1+(1)	• K ⁰ _L	1/2(0-)	• $B^{\pm}/B^{0}/B^{0}_{s}/b$	-baryon AD-	
 ρ(770) 	1+(1)	a2(1700)	$1^{-}(2^{++})$	K_0(800)	$1/2(0^{+})$	MIXTURE		
 ω(782) 	0-(1)	 f₀(1710) 	$0^{+}(0^{++})$	• K*(892) 1/2(1 ⁻)		V _{cb} and V _{ub} CKM Matrix		
 η'(958) 	0+(0-+)	η(1760)	$0^{+}(0^{-+})$	• K1(1270)	$1/2(1^+)$	• B*	$1/2(1^{-})$	
 f₀(980) 	$0^{+}(0^{++})$	 π(1800) 	$1^{-}(0^{-+})$	• K1(1400)	$1/2(1^+)$	B*(5732)	7(7?)	
 a₀(980) 	$1^{-}(0^{++})$	f2(1810)	$0^{+}(2^{++})$	• K*(1410)	$1/2(1^{-})$	e j(erez)	.(.)	
• $\phi(1020)$	$0^{-}(1^{-})$	X(1835)	$?^{?}(?^{-+})$	• $K_{*}^{*}(1430)$ 1/2(0 ⁺)		BOTTOM, STRANGE		
 h1(1170) 	$0^{-(1+-)}$	• \$\phi_3(1850)	$0^{-}(3^{-}-)$	$K^*(1430) = 1/2(2^+)$		$(B = \pm 1, S = \mp 1)$		
• b1(1235)	1+(1+-)	12(1870)	$0^{+}(2^{-+})$	K(1460)	$1/2(2^{-})$	• B ⁰	0(0-)	
• a1(1260)	$1^{-(1++)}$	p(1900)	$1^{+}(1^{-})$	K (1580)	$1/2(0^{-})$	B*	0(1-)	
• fs(1270)	$0^+(2^{++})$	6(1910)	$0^{+}(2^{+})$	K(1630)	1/2(2)	B* (5850)	7(7?)	
• f1(1285)	$0^{+(1++)}$	• fs(1950)	$0^{+}(2^{+}+)$	K (1650)	1/2(1+)	- 35(0000)		
• n(1295)	$0^{+}(0^{-}+)$	03(1990)	1+(3)	$(1(1050)) = 1/2(1^{-1})$		BOTTOM, CHARMED		
 π(1300) 	$1^{-(0^{-+})}$	• fs(2010)	$0^+(2^+)$	- K (1770)	1/2(1)	(<i>B</i> = <i>C</i>	= ±1)	
• a (1320)	$1^{-}(2^{+})$	fo(2020)	$0^{+}(0^{+}+)$	• A2(1770)	1/2(2)	• B_c^{\pm}	0(0-)	
• fo(1370)	$0^{+}(0^{+}+)$	• a (2040)	$1^{-(4^{++})}$	- K (1920)	1/2(3)	-		
h (1380)	$7^{-(1+-1)}$	• fa(2050)	$0^{+}(4^{+}+)$	• K ₂ (1820)	1/2(2)	c	c	
• π ₁ (1400)	$1^{-(1^{-+})}$	$\pi_2(2100)$	$1^{-(2^{-+})}$	K(1830)	1/2(0)	• η _c (15)	0+(0-+)	
• n(1405)	$0^{+}(0^{-}+)$	fo(2100)	$0^{+}(0^{+}^{+})$	R ₀ (1950)	1/2(01)	 J/ψ(15) 	0-(1)	
• £(1420)	$0^{+}(1^{+})$	fo(2150)	$0^{+}(2^{+})$	K ₂ (1980)	1/2(2+)	• $\chi_{c0}(1P)$	0+(0++)	
• (1420)	$0^{-}(1^{-})$	a(2150)	$1^{+}(1^{-})$	 K₄[*](2045) 	1/2(4+)	• $\chi_{c1}(1P)$	$0^+(1^+)$	
fo(1430)	$0^{+}(2^{+}+)$	f.(2200)	$0^{+}(0^{+}+)$	$K_2(2250)$	$1/2(2^{-})$	$h_c(1P)$?!(?!!)	
• 20(1450)	$1^{-}(0^{+}+)$	f.(2220)	$0^+(2 \text{ or } 4^{++})$	K ₃ (2320)	$1/2(3^+)$	• $\chi_{c2}(1P)$	0+(2++)	
• 0(1450)	1+(1)	n(2225)	$0^{+}(0^{-+})$	$K_{5}^{*}(2380)$	1/2(5)	 η_c(25) 	0+(0 - +)	
• n(1475)	$0^{+}(0^{-}+)$	(2250)	1+(3)	K ₄ (2500)	1/2(4-)	• ψ(25)	0-(1)	
• fo(1500)	$0^{+}(0^{+}+)$	• f (2300)	$0^{+}(2^{+}+)$	K(3100)	?!(?!!)	 ψ(3770) 	0-(1)	
£ (1510)	$0^{+}(1^{+}+1)$	£(2300)	$0^{+}(4^{+}+1)$	CHAR	MED	• X(3872)	0'(?'+)	
• f'(1525)	$0^{+}(2^{+}+)$	• fa(2340)	$0^{+}(2^{+}+)$	(()=	+1)	 χ_{c2}(2P) 	$0^+(2^{++})$	
£(1565)	$0^+(2^+)$	0=(2350)	1+(5)	(°	1/0/0-)	Y (3940)	?!(?!!)	
h (1505)	$0^{-(1+-)}$	2.(2450)	$1^{-}(6^{+}+)$	• D ²	1/2(0)	 ψ(4040) 	0-(1)	
m(1595)	$1^{-(1-+)}$	£(2510)	$0^+(6^+)$	• D*	1/2(0)	 ψ(4160) 	0-(1)	
• "1(1000) a (1640)	1 - (1 + +)	<i>i</i> ₆ (2010)	0 (0)	• D*(2007)*	1/2(1)	Y(4260)	??(1)	
£(1640)	$a^{+}(2^{+}^{+})$	OTH	ER LIGHT	• D*(2010)-	1/2(1)	 ψ(4415) 	0-(1)	
12(1040)	$0^{+}(2^{-}+)$	Further States		D ₀ (2400) ⁰	1/2(0+)		T	
• (1650)	$0^{-}(1^{-})$			$D_0^*(2400)^{\pm}$ 1/2(0 ⁺)		bb		
• (1670)	$0^{-}(2^{-})$			• D ₁ (2420) ⁰	$1/2(1^+)$	$\eta_b(1S)$	0+(0-+)	
• a3(10/0)	0 (3)			$D_1(2420)^{\pm}$	1/2(?*)	• T(15)	0-(1)	
				$D_1(2430)^0$	$1/2(1^+)$	• $\chi_{b0}(1P)$	0+(0++)	
				 D[*]₂(2460)⁰ 	1/2(2+)	• $\chi_{b1}(1P)$	$0^+(1^{++})$	
				 D[*]₂(2460)[±] 	$1/2(2^+)$	• χ _{b2} (1P)	0+(2++)	
				D*(2640) [±]	1/2(?')	• T(25)	0-(1)	
				CHARMED	STRANCE	$\Upsilon(1D)$	0-(2)	
				(C=S-	= +1)	 χ_{b0}(2P) 	0+(0++)	
				(c = 5.	0(0=)	• χ _{b1} (2P)	0+(1++)	
				• D _s	0(0)	 χ_{b2}(2P) 	0+(2++)	
				• <i>D</i> ^{*±} _s	0(?:)	• T(35)	0-(1)	
				• $D_{s0}^*(2317)^{\pm}$	0(0+)	• T(45)	0-(1)	
				 D_{\$1}(2460)[±] 	0(1+)	 <i>Υ</i>(10860) 	0-(1)	
				 <i>D</i>_{s1}(2536)[±] 	0(1+)	 <i>Υ</i>(11020) 	0-(1)	
				 D_{s2}(2573)[±] 	0(?')		NDIDATEC	
						NON-99 CA	INDIDATES	

p	P ₁₁	****	∆ (1232)	P ₃₃	****	Λ	P ₀₁	****	Σ+	P ₁₁	****	<i>≡</i> °	P ₁₁	****
n	P ₁₁	****	∆ (1600)	P ₃₃	***	A(1405)	S ₀₁	****	Σ^0	P ₁₁	****	Ξ-	P ₁₁	****
N(1440)	P ₁₁	****	∆ (1620)	S31	****	A(1520)	D ₀₃	****	Σ-	P ₁₁	****	Ξ(1530)	P ₁₃	****
N(1520)	D ₁₃	****	∆ (1700)	D33	****	A(1600)	P ₀₁	***	Σ(1385)	P ₁₃	****	Ξ(1620)		*
N(1535)	S ₁₁	****	∆(1750)	P31	*	A(1670)	S ₀₁	****	Σ(1480)		*	Ξ(1690)		***
N(1650)	S11	****	∆(1900)	S31	**	A(1690)	D ₀₃	****	Σ(1560)		**	Ξ(1820)	D13	***
N(1675)	D15	****	∆(1905)	F35	****	A(1800)	S ₀₁	***	Σ(1580)	D ₁₃	*	Ξ(1950)		***
N(1680)	F ₁₅	****	∆(1910)	P31	****	A(1810)	P ₀₁	***	Σ(1620)	S11	**	Ξ(2030)		***
N(1700)	D ₁₃	***	∆(1920)	P33	***	A(1820)	F ₀₅	****	Σ(1660)	P ₁₁	***	Ξ(2120)		*
N(1710)	P ₁₁	***	∆(1930)	D35	***	A(1830)	D ₀₅	****	Σ(1670)	D ₁₃	****	Ξ(2250)		**
N(1720)	P13	****	∆(1940)	D33	*	A(1890)	P03	****	Σ(1690)		**	Ξ(2370)		**
N(1900)	P13	**	∆(1950)	F37	****	A(2000)		*	Σ(1750)	S ₁₁	***	Ξ (2500)		*
N(1990)	F17	**	∆ (2000)	F35	**	A(2020)	F07	*	Σ(1770)	P ₁₁	*			
N(2000)	F15	**	∆ (2150)	S31	*	A(2100)	G07	****	Σ(1775)	D ₁₅	****	Ω^{-}		****
N(2080)	D13	**	∆ (2200)	G37	*	A(2110)	F ₀₅	***	Σ(1840)	P13	*	$\Omega(2250)^{-}$		***
N(2090)	S11	*	∆(2300)	Hag	**	A(2325)	D ₀₃	*	Σ(1880)	P ₁₁	**	$\Omega(2380)^{-}$		**
N(2100)	P11	*	∆ (2350)	D35	*	A(2350)	Hog	***	Σ(1915)	F15	****	$\Omega(2470)^{-}$		**
N(2190)	G17	****	∆(2390)	F37	*	A(2585)		**	Σ(1940)	D13	***			
N(2200)	D15	**	A(2400)	G39	**	` <i>`</i>			Σ(2000)	S11	*	Λ_c^+		****
N(2220)	H19	****	A(2420)	Ha 11	****				Σ(2030)	F17	****	$\Lambda_{c}(2593)^{+}$		***
N(2250)	G19	****	$\Delta(2750)$	12 12	**				Σ(2070)	F ₁₅	*	$\Lambda_{c}(2625)^{+}$		***
N(2600)	1.11	***	A(2950)	Ka 15	**				Σ(2080)	P13	**	$\Lambda_{c}(2765)^{+}$		*
N(2700)	K1 13	**	(2500)						Σ(2100)	G17	*	$\Lambda_{c}(2880)^{+}$		**
(,	1,10		$\Theta(1540)^+$		*				Σ(2250)		***	$\Sigma_{c}(2455)$		****
									Σ(2455)		**	$\Sigma_{c}(2520)$		***
									Σ(2620)		**	$\Sigma_{c}(2800)$		***
									Σ(3000)		*	Ξ_c^+		***
									Σ(3170)		*	=0		***
			1			1						Ξ'^+		***
												="0		***
		1		$\mathbf{\cap}$								E.(2645)		***
				()								E(2790)		***
												E(2815)		***
				U								Ω_c^0		***
												Ξ_{cc}^+		*
												10		***
												\equiv_b^0, \equiv_b^-		*

This short table gives the name, the quantum numbers (where known), and the status of baryons in the Review. Only the baryons with 3-

or 4-star status are included in the main Baryon Summary Table. Due to insufficient data or uncertain interpretation, the other entries in the short table are not established as baryons. The names with masses are of baryons that decay strongly. For N, Δ , and Ξ resonances, the partial wave is indicated by the symbol $L_{2I,2J}$, where L is the orbital angular momuntum (S, P, D, ...), I is the isospin, and J is the total

**** Existence is certain, and properties are at least fairly well explored.

*** Existence ranges from very likely to certain, but further confirmation is desirable and/or quantum numbers, branching fractions, etc. are not well determined.

** Evidence of existence is only fair.

* Evidence of existence is poor

Wilson non-perturbatively improved Fermions "WORKING HORSE" of lattice QCD calculations

Y. Kuramashi Lattice 2007

Iwasaki gauge action + clover quarks $a^{(-1)} = 2.2 \text{GeV},$ lattice size: $32^3 \times 64$



Heavy lons Collisions and Quark- Gluon Plasma


Phase diagram of QCD



2 dimensional gluodynamics (no dynamical quarks) can be solved analytically

4 dimensional gluodynamics (no dynamical quarks) can be "solved" on laptop



4 dimensional QCD (with dynamical quarks) needs 100-100000 times more CPU time than gluodynamics (we need supercomputers)



u



<u>We do not know how to solve 4 dimensional</u> QCD (with dynamical quarks) at finite chemical potential



u



Phase diagram of QCD



Critical temperature O II

Transition temperature from a variety of studies



- Staggered types, N_f = 2 + 1: p4, asqtad, HISQ, stout — already introduced. Data from only chiral type observables.
- Wilson types, $N_f = 2$:
 - ▷ QCDSF-DIK [arXiv:0910.2392], clover + plaquette, $N_t = 8 - 14$
 - \triangleright WHOT-QCD [arXiv:0909.2121], clover + lwasaki, $N_t = 4 6$
 - \triangleright Brandt et al. [arXiv:1011.6172], clover, $N_t = 16$
 - ▷ tmfT (Florian Burger talk), mtmWilson + treelevel Symazik, $N_t = 8 - 12$
- DWF (HotQCD), $N_f = 2+1$:+lwasaki, $N_t = 8$, $L_s = 32 96$

L. Levkova; Talk at Lattice 2011



AMS 2005, Tampa meeting

created matter at a temperature of about 4 trillion degrees Celsius — the hottest temperature ever reached in a laboratory, about 250,000 times hotter than the center of the Sun

using a giant atom smasher said on they have created a new state of matter - a hot, dense liquid made out of basic atomic particles - and said it shows what the early universe looked like for a very, very brief ime.

"We think we are looking at a phenomenon ... in the universe 13 billion years ago when free quarks and gluons ... cooled down to the particles that we know today," Aronson told a news conference carried by telephone from a meeting of the American Physical Society in Tampa, Fla.

Liquid, not a gas

The quark-gluon plasma was made in the Relativistic Heavy Ion Collider — a powerful atom smasher at Brookhaven National Laboratory in Upton, N.Y. Unexpectedly, the quark-gluon plasma behaved like a perfect liquid of quarks, instead of a gas, the physicists said.

Evidence of 5-th state of matter in heavy ions collisions

- 1. Thermalisation
- 2. Elliptic flow
- 3. Jets quenching
- 4. Spectrum of photons
- 5. Share viscosity eta/s and hydrodynamic approach
- 6. Lattice calculations vs experiment

.....

when they are in *local* thermal equilibrium. Macroscopic currents in one region of the plasma can interact magnetically with other currents in other regions, over tremendous distance scales, creating complicated structures like Fig. 1. Non-Abelian plasmas, however, are somewhat different. From theoretical studies of the equilibrium properties of such plasmas, we know that the non-Abelian interactions cause magnetic *confinement* over distances of order $1/(g^2T)$. It is reasonable to assume that, even dynamically, color magnetic fields cannot exists on distance scales larger than the confinement length. So, unlike traditional electromagnetic plasmas, there are no large-distance magnetic fields. As far as the color degrees of freedom are concerned, the long-distance effective theory of a non-Abelian plasma is hydrodynamics rather than magneto-hydrodynamics.



QUARK-GLUON PLASMA THERMALIZATION AND PLASMA INSTABILITIES PETER ARNOLD

arXiv:hep-ph/0409002v1

Figure 1. Image of a solar coronal filament from NASA's TRACE satellite, from http://antwrp.gsfc.nasa.gov/apod/ap000809.html).

QGP is the thermalized strongly correlated liquid





Collision time is very short and how thermalization occurs it is a question

Below I use a lot of slides made by M.N. Chernodub, P.V. Buividovich and D.E. Kharzeev



Magnetic fields in non-central collisions [Fukushima, Kharzeev, Warringa, McLerran '07-'08]



[1] K. Fukushima, D. E. Kharzeev, and H. J. Warringa, Phys. Rev. D 78, 074033 (2008), URL http://arxiv.org/abs/0808.3382.

[2] D. Kharzeev, R. D. Pisarski, and M. H. G.Tytgat, Phys. Rev. Lett. 81, 512 (1998), URL http://arxiv.org/abs/hep-ph/9804221.

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The medium is filled by electrically charged particles

Large orbital momentum, perpendicular to the reaction plane Large magnetic field along the direction of the orbital momentum



The medium is filled by electrically charged particles

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Comparison of magnetic fields



		D Kharzoov
The Earths magnetic field	0.6 Gauss	D.MIAIZEEV
A common, hand-held magnet	100 Gauss	
The strongest steady magnetic fields achieved so far in the laboratory	4.5 x 10⁵ Ga	uss
The strongest man-made fields ever achieved, if only briefly	10 ⁷ Gauss	
Typical surface, polar magnetic fields of radio pulsars	10 ¹³ Gauss	
Surface field of Magnetars	10 ¹⁵ Gauss	
http://solomon.as.utexas.edu/~duncan/magnetar.html		



Off central Gold-Gold Collisions at 100 GeV per nucleon $e B(\tau = 0.2 \text{ fm}) = 10^3 \sim 10^4 \text{ MeV}^2 \sim 10^{17} \text{ Gauss}$

Magnetic forces are of the order of strong interaction forces

first time in my life I see such effect



We expect the influence of magnetic field on strong interaction physics The effects are nonperturbative, it is impossible to perform analytic calculations and we use

Lattice Calculations



We calculate $\langle \overline{\psi} \Gamma \psi \rangle$; $\Gamma = 1, \gamma_{\mu}, \sigma_{\mu\nu}$

in the external magnetic field and in the presence of the vacuum gluon fields

 \vec{H} external magnetic field



Chiral Magnetic Effect

[Fukushima, Kharzeev, Warringa, McLerran '07-'08]

Electric current appears at regions 1. with non-zero topological charge density 2. exposed to external magnetic field

Experimentally observed at RHIC : charge asymmetry of produced particles at heavy ion collisions

Chiral Magnetic Effect by Fukushima, Kharzeev, Warringa, McLerran



Chiral Magnetic Effect by Fukushima, Kharzeev, Warringa, McLerran 3. Electric current is along magnetic field In the *instanton* field





Chiral Magnetic Effect on the lattice, charge separation

Density of the electric charge vs. magnetic field

H = 0

B=0

$B = (500 \,{ m MeV})^2$



$B = (780 \,{ m MeV})^2$







Chiral Magnetic Effect on the lattice, Non-zero field, subsequent time slices Electric charge density



Chiral Magnetic Effect, EXPERIMENT VS LATTICE DATA (Au+Au)



Preliminary results: conductivity of the vacuum

Qualitative definition of conductivity *****

 $< j_{\mu}(x) j_{\nu}(y) > = C + A \cdot \exp\{-m|x - y|\}$

 $\sigma \propto C$

Preliminary results: conductivity of the vacuum

Conductivity at T>0





Graphene

The Nobel Prize in Physics for 2010 was awarded to Andre Geim and Konstantin Novoselov "for groundbreaking experiments regarding the two-dimensional material graphene"



Featured products

Single Layer Graphene on Copper foil: Graphene Nanopowder: 8 nm Flakes- Q-graphene: 1 gram 4"x2" 5 g



Quantity 1

Buy Now

Quantity 1

Buy No<u>w</u>

Quantity 1

Buy Now

Relativistic particle $E = \sqrt{m^2c^4 + p^2c^2}$ Massless particle E = cp




$$\alpha_{g} = 300\alpha = 2.16 > 1$$

 $\alpha_g > \alpha_g^{crit} = 1.11 \pm 0.06$ Pure graphene is the insulator!



$$\alpha_{g} = 300\alpha = 2.16 > 1$$

 $\alpha_g > \alpha_g^{crit} = 1.11 \pm 0.06$ Pure graphene is the insulator!

If we put graphene on a substrate we can get conductor: α

 $\alpha_g \to \frac{2}{1+\varepsilon} \alpha_g$

Numerical calculation of α_g^{crit}

Is graphene in vacuum an insulator?. Joaquin E. Drut Timo A. Lahde e-Print: **arXiv:0807.0834 [cond-mat.str-el]**, **PRL (2009)**

Monte Carlo Simulation of the Semimetal-Insulator Phase Transition in Monolayer Graphene. W. Armour, Simon Hands, Costas Strouthos Published in **Phys.Rev. B81 (2010) 125105** e-Print: **arXiv:0910.5646 [cond-mat.str-el]**





We can numerically simulate conductor – insulator phase transition!

Magnetic Field and Graphene



Graphene changes its properties when an external magnetic field is applied, we can numerically simulate all that



Along the trajectory of the magnetic head graphene becomes the conductor! We can draw (construct) chips! All that we can simulate on computers

Problems for graphene numerical simulations

Magnetic field Finite temperature Impurities 2-3-4 layers Conductivity Viscosity – Entropy Optical properties Critical indices

Monte Carlo simulation of monolayer graphene at non-zero temperature

Wesley Armour^{a,b}, Simon Hands', and Costas Strouthos^d

arXiv:1105.1043v1 [cond-mat.str-el] 5 May 2011



Graphene has relations with many theoretical problems

Insulator – Conductor



Confinement - Deconfinement



Curvature of the graphene leads to two dimensional gravity for fermions



QCD confinement problems Quark-Gluon plasma in heavy ion collisions Graphene



QCD confinement problems Quark-Gluon plasma in heavy ion collisions Graphene



Graphene